

# Higher Dimensions & Higher Orders in EFT for Gravitational Waves



Ofek Birnholtz



- OB, Shahar Hadar & Barak Kol, Phys. Rev. D 88, 104037 (2013)  
OB & Shahar Hadar, Phys. Rev. D 89, 045003 (2014)  
OB, Shahar Hadar & Barak Kol, IJMPA 29, 24, 30 (2014)  
OB, IJMPA 30, 02, 20 (2015)  
OB & Shahar Hadar, Phys. Rev. D 91, 124065 (2015)

14th Marcel Grossmann Meeting, La Sapienza, Roma  
14 July 2015

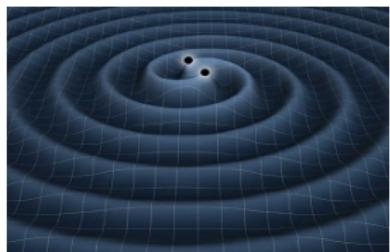
# Outline

- 1 Motivation
- 2 Plan
- 3 Math
- 4 Results
- 5 Higher Order interactions
- 6 Summary & Future

# Outline

- 1 Motivation
- 2 Plan
- 3 Math
- 4 Results
- 5 Higher Order interactions
- 6 Summary & Future

## Gravitational Waves from a 2-body system



$$S = \frac{1}{16\pi G_d} \int \sqrt{-g} R d^d x - \frac{1}{2} \int h_{\mu\nu} T^{\mu\nu} d^d x,$$

$$\nabla_\mu T^{\mu\nu} = 0 , \quad g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$

(  $c = 1$ , linearized source )

Issues:

- Metric fields are numerous, mixed, of different behaviours
- Some degrees of freedom are pure gauge
- Number of d.o.f (real and gauge) change with dimension
- Radiation and system zones have different symmetries
- Problem part-conservative, part dissipative
- Solution sensitive to new effects & increasing order

# Outline

1 Motivation

2 Plan

3 Math

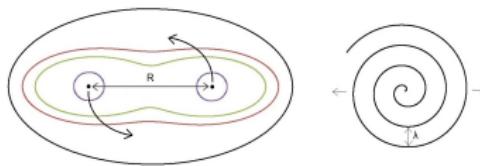
4 Results

5 Higher Order interactions

6 Summary & Future

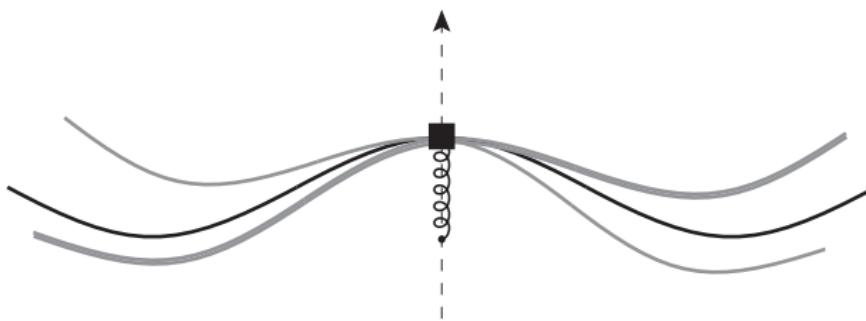
“Everything should be made as simple as possible,  
but not simpler” - Albert Einstein

- Zone Separation ( $\lambda \propto \frac{R}{v} \gg R$ )
  - System Zone:  $\sim$  stationarity, NRG fields
  - Radiation Zone:  $\sim$  (hyper-)spherical symmetry



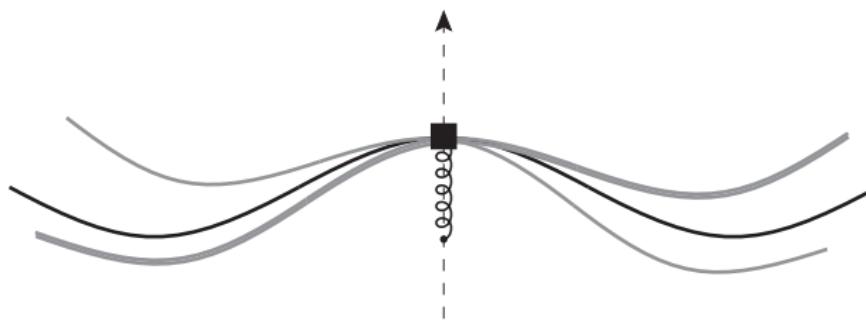
- Rad. Z. : Dimensional Reduction + Gauge Invariance  
 $\implies$  Master-Fields  $\hbar_S, \hbar_V, \hbar_T$
- Sys. Z. : Match Master Sources  $\mathcal{T}_S, \mathcal{T}_V, \mathcal{T}_T$   
 $\implies$  Master Action & Master Equation for outgoing waves

“Everything should be made as simple as possible,  
but not simpler” - Albert Einstein



- Rad. Z. : Dimensional Reduction + Gauge Invariance  
 $\implies$  Master-Fields  $\hbar_S, \hbar_V, \hbar_T$
- Sys. Z. : Match Master Sources  $\mathcal{T}_S, \mathcal{T}_V, \mathcal{T}_T$   
 $\implies$  Master Action & Master Equation for outgoing waves

“Everything should be made as simple as possible,  
but not simpler” - Albert Einstein



- Rad. Z. : Dimensional Reduction + Gauge Invariance  
 $\implies$  Master-Fields  $\hbar_S, \hbar_V, \hbar_T$
- Sys. Z. : Match Master Sources  $\mathcal{T}_S, \mathcal{T}_V, \mathcal{T}_T$   
 $\implies$  Master Action & Master Equation for outgoing waves  
 $\implies$  Field Doubling for Radiation-Reaction Effective Action

# Outline

- 1 Motivation
- 2 Plan
- 3 Math
- 4 Results
- 5 Higher Order interactions
- 6 Summary & Future

## Finding the G.I. Master Fields

Starting from

$$h_{\alpha\beta} = \sum \begin{pmatrix} h_{tt} n_L & h_{tr} n_L & h_t \partial_\Omega n_L + h_{tV_N} n_{V_N}^L \\ \cdots & h_{rr} n_L & h_r \partial_\Omega n_L + h_{rV_N} n_{V_N}^L \\ \cdots & \cdots & h_S n_{\Omega\Omega'}^L + \tilde{h}_S \tilde{n}_{\Omega\Omega'}^L + h_{V_N} n_{V_N\Omega\Omega'}^L + h_{\nabla\nabla} n_{\nabla\nabla\Omega\Omega'}^L \end{pmatrix} e^{-i\omega t}$$

we find the tensors Master fields

$$\tilde{h}_{\nabla\nabla} = r^{-(\ell+2)} h_{\nabla\nabla}.$$

Using 1 algebraic equation we find the vector Master fields

$$\tilde{h}_{V_N} = -\frac{\ell(\ell+\hat{d}+1)}{4r^\ell} \left( \omega^2 - \frac{\hat{c}_s}{r^2} \right)^{-1} \left( \frac{2h_{tV_N} + i\omega h_{V_N}}{r^2} \right)'$$

Using 3 algebraic equations we find the scalars  $\Phi := h_{tt} + 2i\omega h_t - \omega^2 \tilde{h}_S$ , canonically transformed to  $\tilde{\Phi}$  with  $F[\Phi, \Phi^*, \tilde{\Phi}, \tilde{\Phi}^*] = -\hat{c}_s r^{\frac{\hat{d}-1}{2}} (\tilde{\Phi}^* \Phi + \tilde{\Phi} \Phi^*)$ , and finally the scalar Master fields

$$\tilde{h}_S = -\frac{\hat{d}(\ell+\hat{d}+1)}{2\ell} r^{-(\ell+\frac{\hat{d}+1}{2})} \tilde{\Phi}$$

# 1D Master Action & Master Equation

Master Action:

$$S^\epsilon = \frac{1}{2} \sum_{L\omega} \int dr \left[ \frac{N_{\ell,\hat{d}}}{G_d R_{\ell,\hat{d}}^\epsilon} r^{2\ell+\hat{d}+1} \mathfrak{h}_\epsilon^* \mathfrak{L} \mathfrak{h}^\epsilon - (\mathfrak{h}_\epsilon^* \mathcal{T}^\epsilon + c.c.) \right]$$

with

$$N_{\ell,\hat{d}} = \frac{\Gamma(1+\hat{d}/2)}{2^\ell \Gamma(1+\alpha)} = \frac{\hat{d}!!}{(2\ell+\hat{d})!!} , \quad \alpha = \ell + \frac{\hat{d}}{2} , \quad M_{\ell,\hat{d}} = \frac{\pi}{2^{2\alpha+1} N_{\ell,\hat{d}} \Gamma^2(\alpha+1)}$$

$$R_{\ell,\hat{d}}^S = \frac{\hat{d}(\ell+\hat{d}+1)(\ell+\hat{d})}{(\ell-1)\ell} , \quad R_{\ell,\hat{d}}^V = \frac{(\hat{d}+1)(\ell+\hat{d}+1)\ell}{2(\ell-1)(\ell+\hat{d})} , \quad R_{\ell,\hat{d}}^T = \frac{8(\hat{d}+1)}{\hat{d}^2 c_s (c_s - \hat{d})}$$

Master Equation:

$$0 = \frac{\delta S}{\delta \mathfrak{h}_\epsilon^{L\omega*}} = \frac{N_{\ell,\hat{d}}}{G_d R_{\ell,\hat{d}}^\epsilon} r^{2\ell+\hat{d}+1} \left( \omega^2 + \partial_r^2 + \frac{2\ell+\hat{d}+1}{r} \partial_r \right) \mathfrak{h}_{L\omega}^\epsilon - \mathcal{T}_{L\omega}^\epsilon$$

## Field doubling

Double the field and the source (Schwinger, Galley) :

- These reflect directed propagation
- Interpretation: "pulling" mirror / radiation "sink"

$$\hat{h} \rightarrow \hat{h}, \quad , \quad Q \rightarrow \hat{Q} = \frac{\delta Q}{\delta x} \hat{x}$$

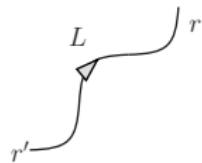


Doubled action:

$$\begin{aligned} \hat{S} \left[ h, \hat{h}; Q, \hat{Q} \right] &= \int \left[ \frac{\delta S}{\delta h} \hat{h} + \frac{\delta S}{\delta Q} \hat{Q} \right] \\ &= \int \hat{h}_\epsilon^* \left[ \frac{N_{\ell, \hat{d}}}{G_d R_{\ell, \hat{d}}^\epsilon} r^{2\ell + \hat{d} + 1} \left( \omega^2 + \partial_r^2 + \frac{2\ell + \hat{d} + 1}{r} \partial_r \right) h^\epsilon - Q_\epsilon \right] \end{aligned}$$

EOM found by varying w.r.t.  $\hat{h}$

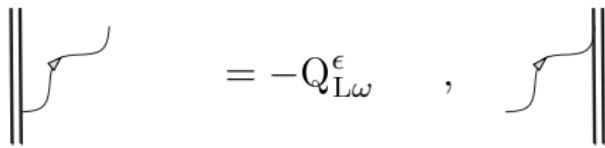
## Feynman Rules



$$= G_{\text{ret}}^\epsilon(r', r) = -i G \omega^{2\ell+\hat{d}} M_{\ell,\hat{d}} R_{\ell,\hat{d}}^\epsilon \tilde{j}_\alpha(\omega r_<) \tilde{h}_\alpha^+(\omega r_>) \delta_{LL'}$$

$$\tilde{j}_\alpha := \Gamma(\alpha + 1) 2^\alpha \frac{j_\alpha(x)}{x^\alpha} = \sum_{p=0}^{\infty} \frac{(-)^p (2\alpha)!!}{(2p)!! (2p+2\alpha)!!} x^{2p} = 1 - \frac{x^2}{2(2\alpha+2)} + \dots$$

$$\tilde{h}_\alpha := \Gamma(\alpha + 1) 2^\alpha \frac{h_\alpha(x)}{x^\alpha}$$



$$= -Q_{L\omega}^\epsilon \quad , \quad = -\hat{Q}_{L\omega}^{\epsilon*}$$

$$\hat{S} = \left\| \begin{array}{c} \text{loop} \\ \text{vertical line} \end{array} \right\| = \sum \left( \hat{Q} G Q \right) \quad , \quad F = \frac{\delta \hat{S}}{\delta \hat{x}}$$

# Outline

- 1 Motivation
- 2 Plan
- 3 Math
- 4 Results
- 5 Higher Order interactions
- 6 Summary & Future

## Quadrupole Moments (LO, NLO)

- $Q_E^{ij}$  : Mass quadrupole

$$Q_S^{ij} \rightarrow \sum_{A=1}^n m_A \left( x^i x^j - \frac{1}{D} \delta^{ij} x^2 \right)_A$$

- $Q_M^{ij}$  : Current quadrupole

$$Q_M^{L_2} = Q_M^{ij} = 2 \sum_{A=1}^n \left[ m \left( \vec{r} \wedge \vec{J} \right)^{(i} x^{j)} \right]_A$$

- $Q_E^{ijk}$  : Mass octupole

$$Q_S^{L_3} = Q_S^{ijk} = \sum_{A=1}^n m_A \left[ x^i x^j x^k - \frac{1}{D+2} (\delta^{ij} x^k + \delta^{ik} x^j + \delta^{jk} x^i) x^2 \right]_A$$

## General Dimension - RR Effective Action

$$\hat{S}_{\text{linear}} = \left\| h_S \right\| + \left\| h_M \right\| + \left\| h_T \right\|$$

## General Dimension - RR Effective Action

$$\begin{aligned}
 \hat{S}_{\text{linear}} &= \left\| h_S \right\| + \left\| h_M \right\| + \left\| h_T \right\| \\
 &= \int dt \sum_{\ell} \frac{G_d (-)^{\ell+\hat{d}} (\ell+\hat{d}+1)}{\hat{d}!! (2\ell+\hat{d})!! (\ell-1)} \left[ \frac{\hat{d}(\ell+\hat{d})}{\ell} \hat{Q}_{(E)}^L \partial_t^{2\ell+\hat{d}} Q_L^{(E)} \right. \\
 &\quad \left. + \frac{(\hat{d}+1)\ell}{2(\ell+\hat{d})} \hat{Q}_{(M)}^L \partial_t^{2\ell+\hat{d}} Q_L^{(M)} + \#_{(T)} \right]
 \end{aligned}$$

## General Dimension - RR Effective Action

$$\begin{aligned}
 \hat{S}_{\text{linear}} &= \left\| h_S \right\| + \left\| h_M \right\| + \left\| h_T \right\| \\
 &= \int dt \sum_{\ell} \frac{G_d (-)^{\ell+\hat{d}} (\ell+\hat{d}+1)}{\hat{d}!! (2\ell+\hat{d})!! (\ell-1)} \left[ \frac{\hat{d}(\ell+\hat{d})}{\ell} \hat{Q}_{(E)}^L \partial_t^{2\ell+\hat{d}} Q_L^{(E)} \right. \\
 &\quad \left. + \frac{(\hat{d}+1)\ell}{2(\ell+\hat{d})} \hat{Q}_{(M)}^L \partial_t^{2\ell+\hat{d}} Q_L^{(M)} + \#_{(T)} \right]
 \end{aligned}$$

$$\hat{S}_{\text{LO+NLO}}^{(d=4)} = G_d \int dt \left[ -\frac{1}{5} \hat{Q}_E^{ij} \partial_t^5 Q_E^{ij} - \frac{4}{45} \hat{Q}_M^{ij} \partial_t^5 Q_M^{ij} + \frac{1}{189} \hat{Q}_E^{ijk} \partial_t^7 Q_E^{ijk} \right]$$

## General d Gravitation RR Effective Action

$$S_{\text{eff}}^{\text{LO}} = (-)^{\frac{\hat{d}+1}{2}} G_d \int dt \frac{\hat{d}(\hat{d}+2)(\hat{d}+3)}{2 \hat{d}!! (\hat{d}+4)!!} \hat{Q}_S^{L_2} \partial_t^{d+1} Q_S^{L_2}$$

$$\begin{aligned} S_{\text{eff}}^{\text{NLO}} = & (-)^{\frac{\hat{d}+1}{2}} G_d \int dt \left[ \frac{\hat{d}(\hat{d}+2)(\hat{d}+3)}{2 \hat{d}!! (\hat{d}+4)!!} \left( \hat{Q}_S^{L_2} \partial_t^{d+1} \delta^1 Q_S^{L_2} + \delta^1 \hat{Q}_S^{L_2} \partial_t^{d+1} Q_S^{L_2} \right) \right. \\ & - \frac{\hat{d}(\hat{d}+4)(\hat{d}+3)}{6 \hat{d}!! (\hat{d}+6)!!} \hat{Q}_S^{L_3} \partial_t^{d+3} Q_S^{L_3} \\ & \left. + \frac{2 \hat{d} (\hat{d}+3)}{3 \hat{d}!! (\hat{d}+4)!!} \hat{Q}_M^{L_2} \partial_t^{d+1} Q_M^{L_2} \right]. \end{aligned}$$

# Outline

- 1 Motivation
- 2 Plan
- 3 Math
- 4 Results
- 5 Higher Order interactions
- 6 Summary & Future

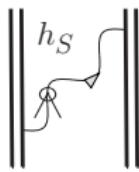
## In System Zone: Quadrupole Moments (LO, NLO)

- $\delta^1 Q_E^{ij}$  : Mass quadrupole (+1PN corrected including first system zone nonlinear effect: gravitating potential energy,  $\sim -\frac{Gm_A m_B}{r^{\hat{d}}}$ )

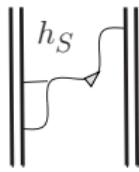


$$\begin{aligned} \delta^1 Q_S^{L_2} = & \sum_{A=1}^n m_A \left[ \left( \frac{\hat{d}+2}{2\hat{d}} v_A^2 - \sum_{B \neq A} \frac{G_d m_B}{\|\vec{x}_A - \vec{x}_B\|^{\hat{d}}} \right) x_A^{L_2} \right. \\ & - \frac{2(\hat{d}+1)}{\hat{d}(\hat{d}+2)} \partial_t \left( \vec{x}_A \cdot \vec{v}_A x_A^{L_2} \right) \\ & \left. + \frac{(\hat{d}^2+6\hat{d}+4)}{2\hat{d}(\hat{d}+2)(\hat{d}+6)} \partial_t^2 \left( r_A^2 x_A^{L_2} \right) \right] \end{aligned}$$

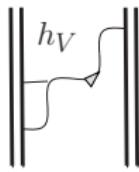
## In Radiation Zone: Background Interactions



Scalar / Cosmological Constant  $\Lambda$



Scalar / Total Mass Curvature  $(+(1 + \frac{\hat{d}}{2})\text{PN})$



Vector / Total Mass Curvature  $(+(2 + \frac{\hat{d}}{2})\text{PN})$

## Spin Effects

co-rotation    maximal-spin

---

$$\text{Spin-Orbit} \quad 1 + 2/\hat{d} \quad 1/2 + 1/\hat{d}$$

$$\text{Spin-Spin} \quad 4/\hat{d} \quad 2/\hat{d}$$

# Outline

- 1 Motivation
- 2 Plan
- 3 Math
- 4 Results
- 5 Higher Order interactions
- 6 Summary & Future

## GR radiation-reaction in general dimension

- Leading order and Next-to-leading order in any dimension
- Joint analytical Action formulation of radiation & reaction
- Economization of traditional computations
- Nonlinear effects - separately for Sys. and Rad. zones
- Beginnings of +1.5PN, +2PN in any dimension
- Ready platform for other effects

## Plans for the future

- The 2-body problem for systems of compact objects:
  - Complete 4PN, 4.5PN, 5PN radiative effects
  - Use EFT to manage Spin and tidal effects
  - Study scattering trajectories in 4d and higher
  - Numerical recipes from EFT
  - Parameter determination from GW signals
- Other high-D systems
  - Explore hyper-relativistic EM problem (Cerenkov)
  - Explore ultra-relativistic systems
- Develop Analytical basis for Effective One Body

Thank you for your attention



Questions?

# Image credits



K. Thorne (Caltech) & T. Carnahan (NASA GSFC)



J. Tenniel in L. Carroll, "Through the Looking-Glass, and What Alice Found There" (1871)



Flickr's [Flood G](#), (creative commons license)