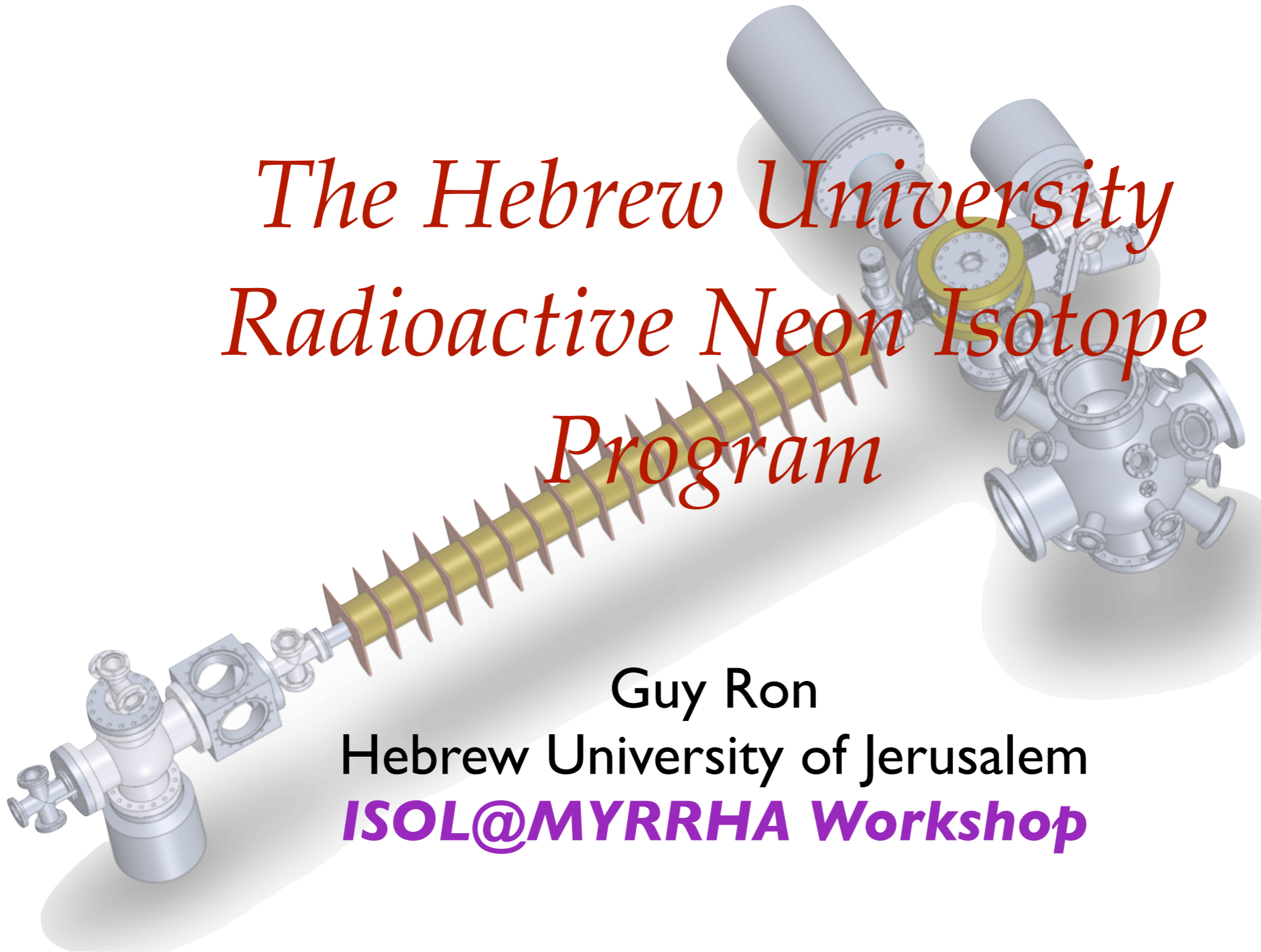




# *The Hebrew University Radioactive Neon Isotope Program*



Guy Ron

Hebrew University of Jerusalem

***ISOL@MYRRHA Workshop***

# Outline

- (Very) short intro
- Why neon?
- Some technical aspects of the HUIJI program
- Production schemes and relevance to MYRRHA.

# WHERE TO SEARCH FOR BEYOND SM PHYSICS?

- ✿ Brute force (*“Swifter, Higher, Stronger”*):
  - ✿ Go higher in energy/luminosity.
  - ✿ LHC/Tevatron/ILC.
- ✿ Finesse:
  - ✿ High precision experiments.
  - ✿ Detect the effect of beyond SM on low energy observables.
  - ✿ “Table top” experiments:  $0\nu\beta\beta$ , atomic PNC, EDM,  $\nu\beta$  correlation.
  - ✿ Accelerator based: Proton/Neutron weak charge ( $Q_{\text{weak}}$ ,  $PR_{\text{ex}}$ , ...).

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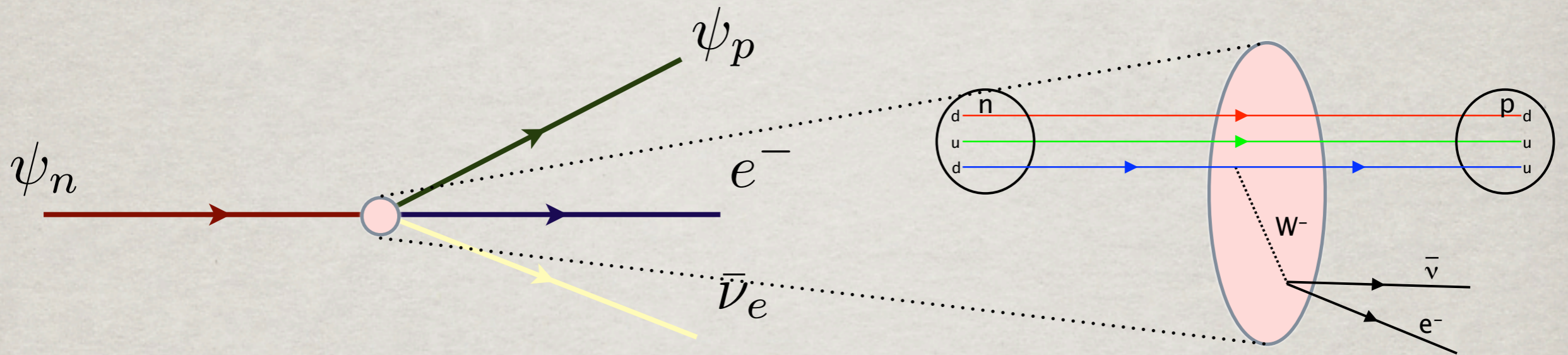
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*Weak Interaction*

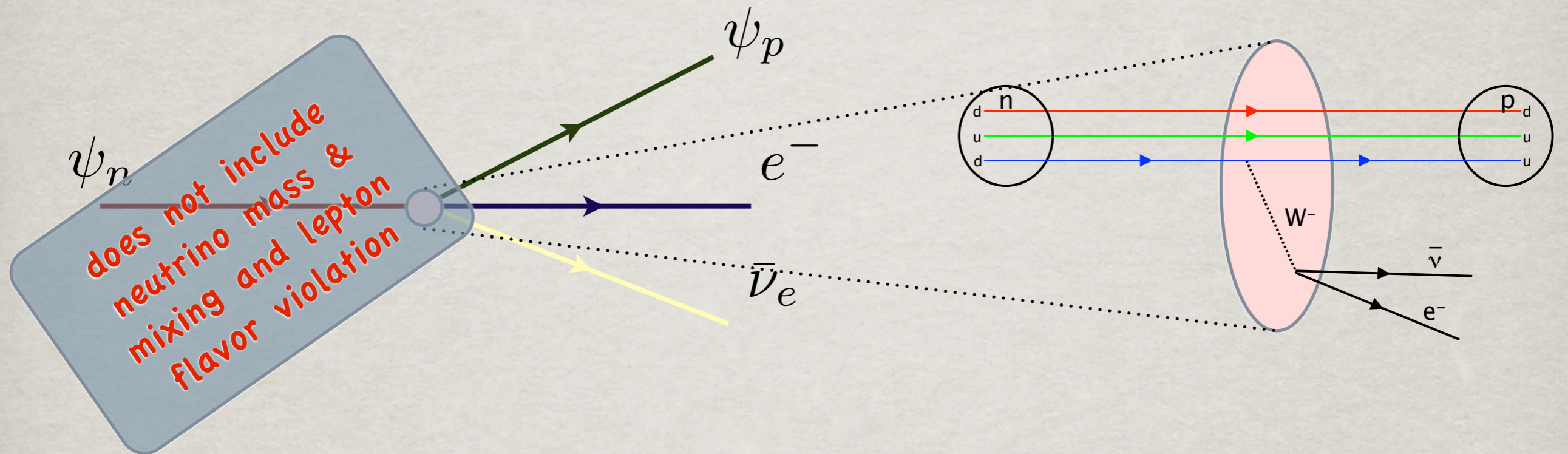
# NUCLEAR $\beta$ DECAY



$$\begin{aligned}
 H_\beta = & (\bar{\psi}_n \psi_p) (C_s \bar{\psi}_e \psi_\nu + C'_s \bar{\psi}_e \psi_\nu \gamma_5 \psi_\nu) \\
 & + (\bar{\psi}_n \gamma_\mu \psi_p) (C_V \bar{\psi}_e \gamma^\mu \psi_\nu + C'_V \bar{\psi}_e \gamma^\mu \gamma_5 \psi_\nu) \\
 & + \frac{1}{2} (\bar{\psi}_n \sigma_{\lambda\nu} \psi_p) (C_T \bar{\psi}_e \sigma^{\lambda\nu} \psi_\nu + C'_T \bar{\psi}_e \sigma^{\lambda\nu} \gamma_5 \psi_\nu) \\
 & - (\bar{\psi}_n \gamma_\mu \gamma_5 \psi_p) (C_A \bar{\psi}_e \gamma^\mu \gamma_5 \psi_\nu + C'_A \bar{\psi}_e \gamma^\mu \psi_\nu) \\
 & + (\bar{\psi}_n \gamma_5 \psi_p) (C_P \bar{\psi}_e \gamma_5 \psi_\nu + C'_P \bar{\psi}_e \psi_\nu)
 \end{aligned}$$

19 free parameters  
(10 complex couplings  
- arbitrary phase)

# NUCLEAR $\beta$ DECAY

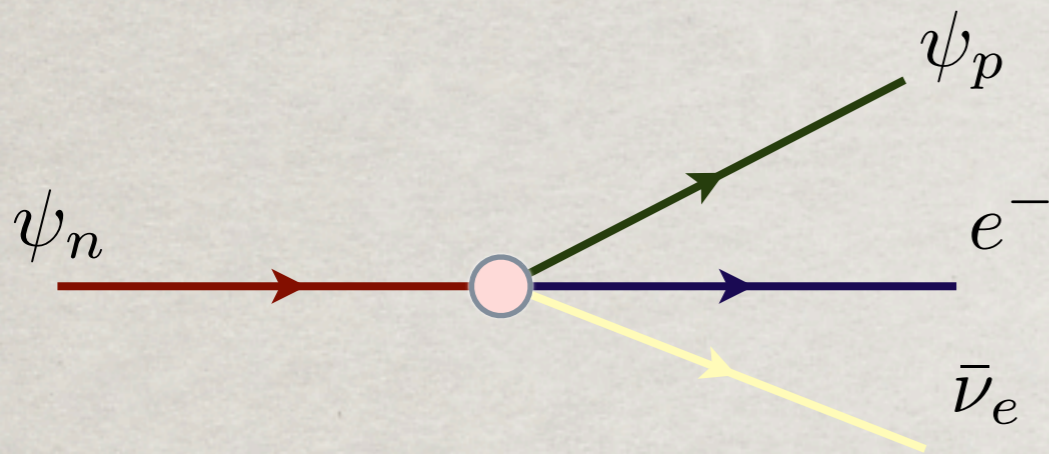


does not include  
neutrino mass &  
mixing and lepton  
flavor violation

$$\begin{aligned}
 H_\beta = & (\bar{\psi}_n \psi_p) (C_s \bar{\psi}_e \psi_\nu + C'_s \bar{\psi}_e \psi_\nu \gamma_5 \psi_\nu) \\
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 \end{aligned}$$

19 free parameters  
(10 complex couplings  
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This is Standard Model

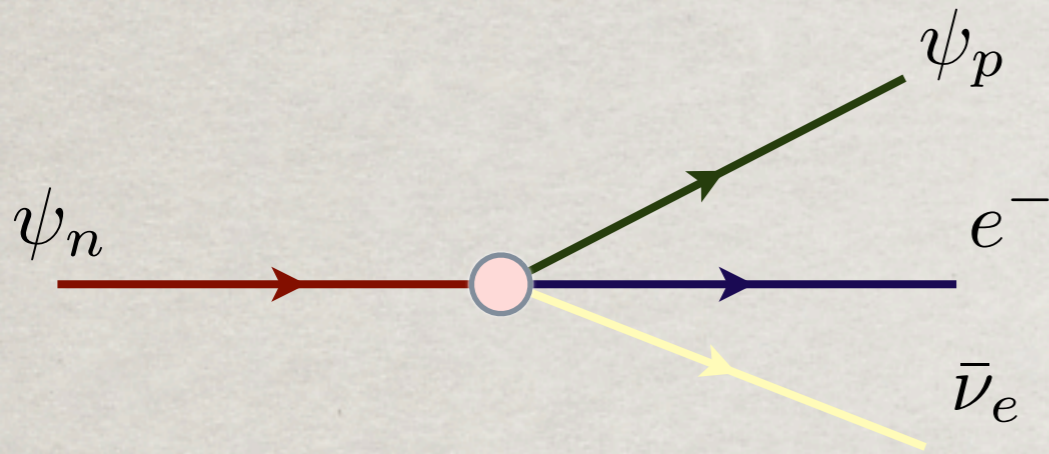


$$H_\beta = (\bar{\psi}_n \gamma_\mu \psi_p) (C_V \bar{\psi}_e \gamma^\mu \psi_\nu + C'_V \bar{\psi}_e \gamma^\mu \gamma_5 \psi_\nu) - (\bar{\psi}_n \gamma_\mu \gamma_5 \psi_p) (C_A \bar{\psi}_e \gamma^\mu \gamma_5 \psi_\nu + C'_A \bar{\psi}_e \gamma^\mu \psi_\nu)$$

$$C_V = C'_V = 1$$

$$C_A = C'_A = 1.26$$

# This is Standard Model

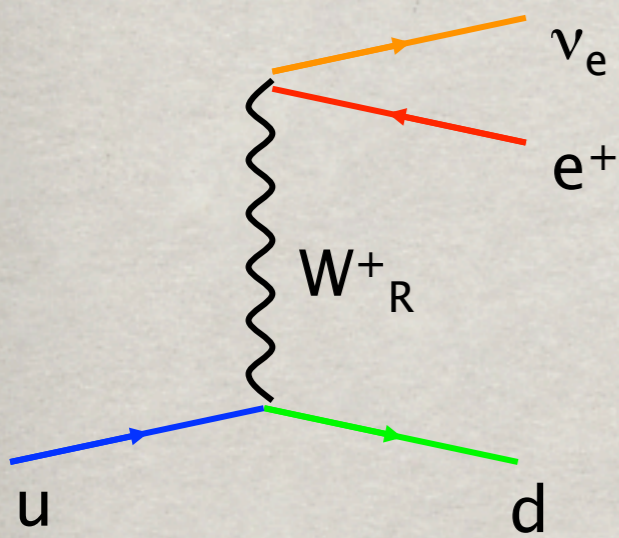


$$H_\beta = (\bar{\psi}_n \gamma_\mu \psi_p) (C_V \bar{\psi}_e \gamma^\mu \psi_\nu + C'_V \bar{\psi}_e \gamma^\mu \gamma_5 \psi_\nu) - (\bar{\psi}_n \gamma_\mu \gamma_5 \psi_p) (C_A \bar{\psi}_e \gamma^\mu \gamma_5 \psi_\nu + C'_A \bar{\psi}_e \gamma^\mu \psi_\nu)$$

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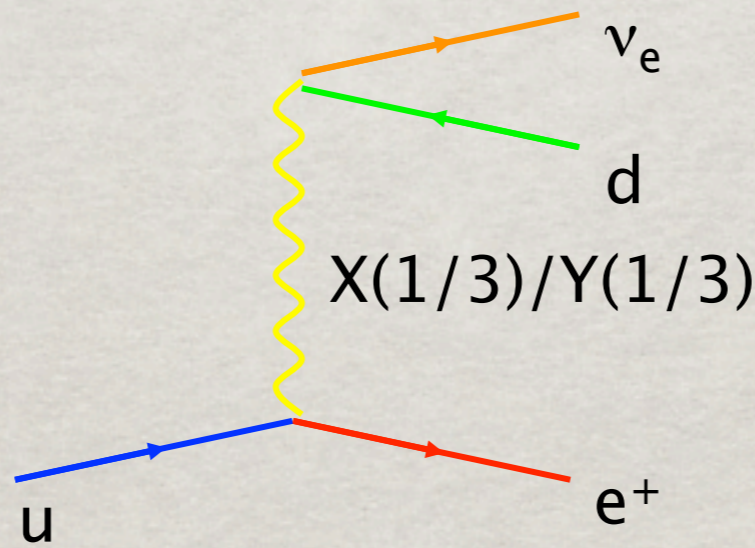
$$C_A = C'_A = 1.26$$

# This is Not....



Right handed bosons

$$C \neq C'$$



Scalar or Tensor  
Leptoquarks

$$C_T \neq 0$$

$$C_S \neq 0$$

- SUSY slepton flavor mixing.
- SUSY LR mixing.
- many more (with different C's)...



# $\beta$ DECAY 101

Total decay rate (electron polarization not detected)

$$\frac{d\Gamma}{dE_\beta d\Omega_\beta d\Omega_\nu} \propto \xi \left\{ 1 + a \frac{\vec{p}_e \cdot \vec{p}_\nu}{E_e E_\nu} + b \frac{m}{E_e} + c \left[ \frac{1}{3} \frac{\vec{p}_e \cdot \vec{p}_\nu}{E_e E_\nu} - \frac{(\vec{p}_e \cdot \vec{j})(\vec{p}_\nu \cdot \vec{j})}{E_e E_\nu} \right] \right. \\ \left. \left[ \frac{J(J+1) - 3 \langle (\vec{J} \cdot \vec{j})^2 \rangle}{J(2J-1)} \right] + \frac{\langle \vec{J} \rangle}{J} \cdot \left[ A \frac{\vec{p}_e}{E_e} + B \frac{\vec{p}_\nu}{E_\nu} + D \frac{\vec{p}_e \times \vec{p}_\nu}{E_e E_\nu} \right] \right\}$$

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Electron-neutrino correlation

$$\xi a = |M_F|^2 \left( -|C_S|^2 + |C_V|^2 - |C'_S|^2 + |C'_V|^2 \right) + \\ \frac{|M_{GT}|^2}{3} \left( |C_T|^2 - |C_A|^2 + |C'_T|^2 - |C'_A|^2 \right) \\ \xi = |M_F|^2 \left( |C_S|^2 + |C_V|^2 + |C'_S|^2 + |C'_V|^2 \right) + \\ |M_{GT}|^2 \left( |C_T|^2 + |C_A|^2 + |C'_T|^2 + |C'_A|^2 \right)$$

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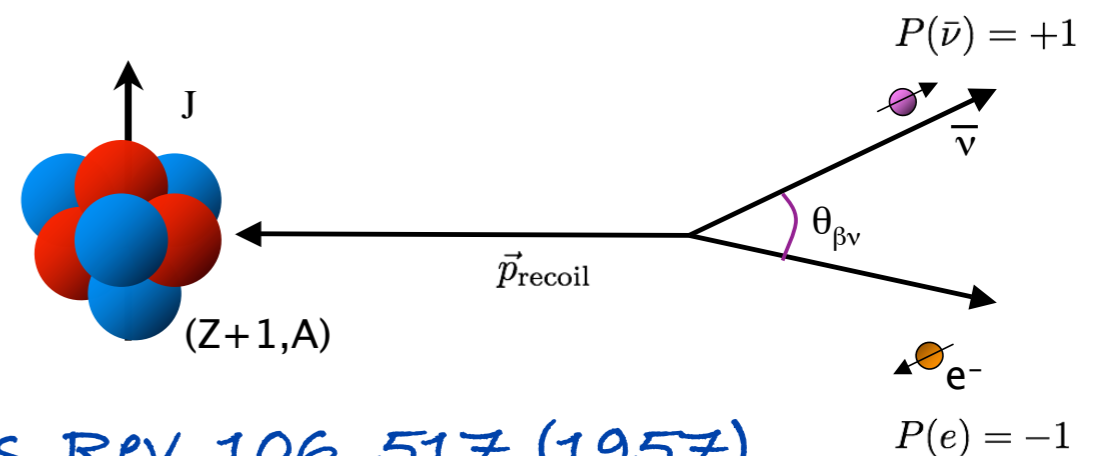
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$\beta + \nu$  carry no AM  $\rightarrow$  emitted in same direction (opposite helicities)

Pure Fermi:  $a = 1$



# $\beta$ DECAY 101

Total decay rate (electron polarization not detected)

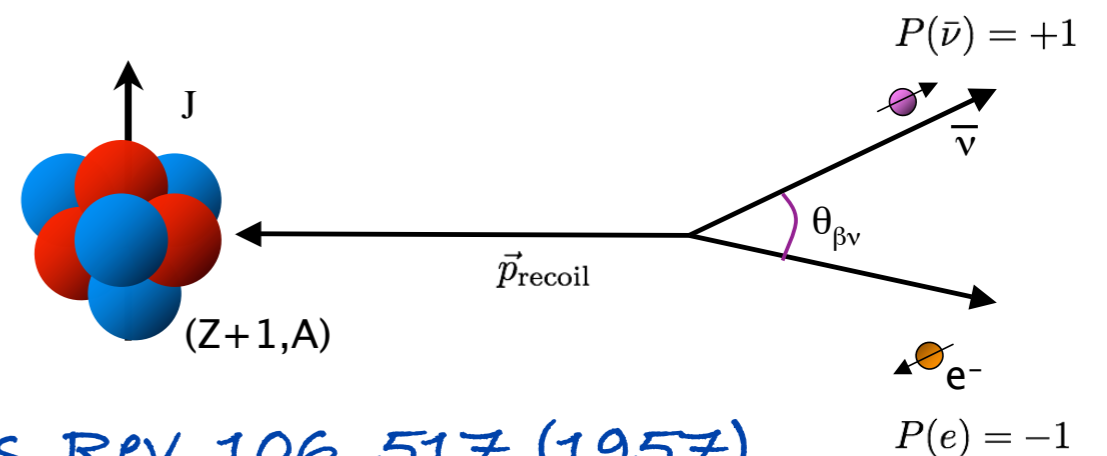
$$\frac{d\Gamma}{dE_\beta d\Omega_\beta d\Omega_\nu} \propto \xi \left\{ 1 + a \frac{\vec{p}_e \cdot \vec{p}_\nu}{E_e E_\nu} + b \frac{m}{E_e} + c \left[ \frac{1}{3} \frac{\vec{p}_e \cdot \vec{p}_\nu}{E_e E_\nu} - \frac{(\vec{p}_e \cdot \vec{j})(\vec{p}_\nu \cdot \vec{j})}{E_e E_\nu} \right] \right. \\ \left. \left[ \frac{J(J+1) - 3 \langle (\vec{J} \cdot \vec{j})^2 \rangle}{J(2J-1)} \right] + \frac{\langle \vec{J} \rangle}{J} \cdot \left[ A \frac{\vec{p}_e}{E_e} + B \frac{\vec{p}_\nu}{E_\nu} + D \frac{\vec{p}_e \times \vec{p}_\nu}{E_e E_\nu} \right] \right\}$$

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**$\beta + \nu$  carry 1 unit AM  $\rightarrow$  emitted in opposite directions (factor of 3 from spin directions)**

Pure GT:  $a = -1/3$



# β DECAY 101

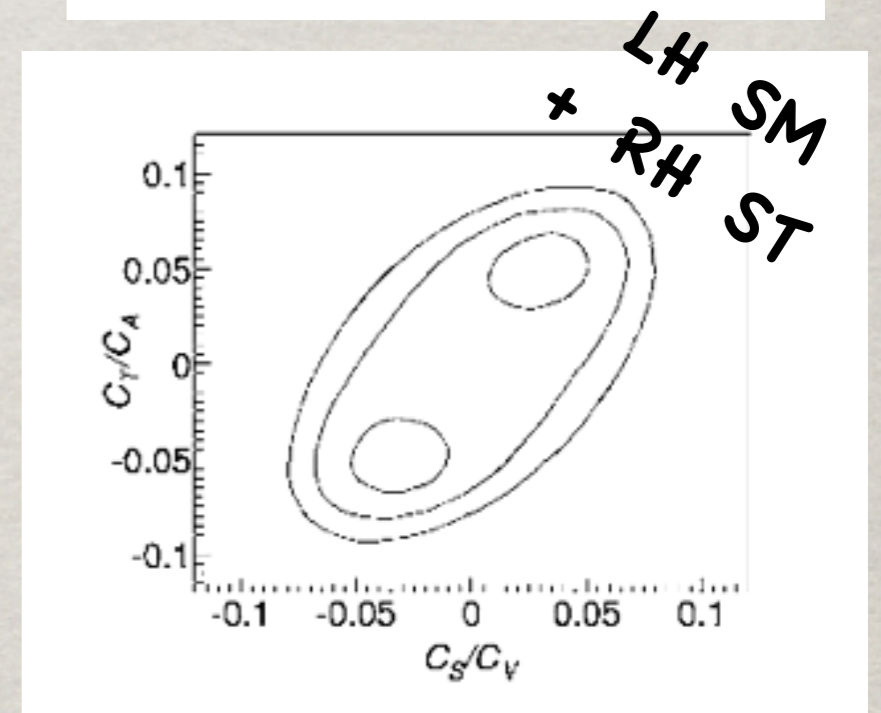
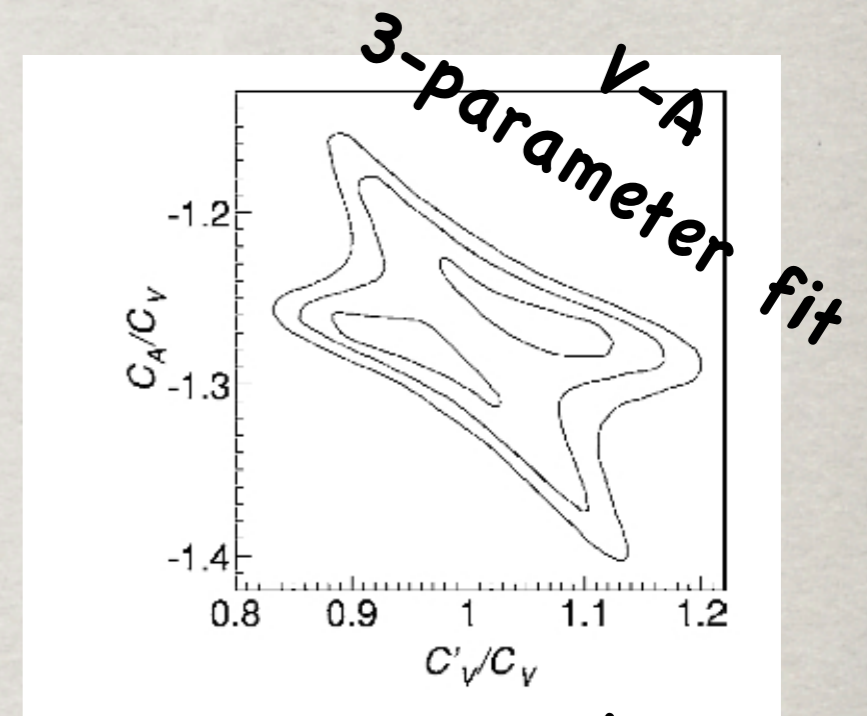
Possible observables in nuclei

$$\frac{d\Gamma}{dE_\beta d\Omega_\beta d\Omega_\nu} \propto \xi \left\{ 1 + a \frac{\vec{p}_e \cdot \vec{p}_\nu}{E_e E_\nu} + b \frac{m}{E_e} + c \left[ \frac{1}{3} \frac{\vec{p}_e \cdot \vec{p}_\nu}{E_e E_\nu} - \frac{(\vec{p}_e \cdot \vec{j})(\vec{p}_\nu \cdot \vec{j})}{E_e E_\nu} \right] \right. \\ \left. \left[ \frac{J(J+1) - 3 \langle (\vec{J} \cdot \vec{j})^2 \rangle}{J(2J-1)} \right] + \frac{\langle \vec{J} \rangle}{J} \cdot \left[ A \frac{\vec{p}_e}{E_e} + B \frac{\vec{p}_\nu}{E_\nu} + D \frac{\vec{p}_e \times \vec{p}_\nu}{E_e E_\nu} \right] \right\}$$

Parameter	Observable	Sensitivity	SM Prediction
<b>a</b>	<b>β-ν (recoil) correlation</b>	<b>Tensor &amp; Scalar terms</b>	<b>1 for pure Fermi -1/3 for pure GT or combination</b>
<b>b</b> (Fierz term)	<b>Comparison of β<sup>+</sup> to EC rate</b>	<b>SV/T/A interference</b>	<b>0</b>
<b>A</b>	<b>β asymmetry for polarized nuclei</b>	<b>Tensor, ST/VA Parity</b>	<b>Nucleus dependent</b>
<b>B</b>	<b>ν asymmetry (recoil) for polarized nuclei</b>	<b>Tensor, TA/ST/VA/SA/VT Parity</b>	<b>Nucleus dependent</b>
<b>D</b>	<b>Triple product</b>	<b>ST/VA Interference TRI</b>	<b>0</b>

# LIMITS ON NON-SM COUPLING

- ✱ Very large model space.
- ✱ Not spanned by collider experiments.
- ✱ Current best limits not very stringent.
- ✱ Naively  $\frac{C_T}{C_A}, \frac{C_S}{C_V} \propto \left( \frac{M_W}{M_{NewPhys}} \right)^2$   
so uncertainty to 0.01 probes new physics at  $\sim 1\text{TeV}$ !
- ✱ Possible effects on high energy results (W production at D0).



*N. Severijns, M. Beck, and O. Naviliat-Cuncic, Rev. Mod. Phys. 78, 991 (2006)*  
*J. Sromki, AIP Conf Proc 338 (1995)*

# ANATOMY OF AN EXPERIMENT

**Produce Radioactive Atoms**

*(Produce, Transport, Neutralize)*



**Trap**

*(MOT, Dipole, Ion, Electrostatic)*



**Wait...**



**Detect decay products ( $\beta$ , Ion)**

*(Scintillators, MCPs,...)*



**Analyze and compare to SM**

# Why Neon? - (I) The Period Table

1 <b>H</b> Hydrogen 1																	2 <b>He</b> Helium 4
3 <b>Li</b> Lithium 7	4 <b>Be</b> Beryllium 9											5 <b>B</b> Boron 11	6 <b>C</b> Carbon 12	7 <b>N</b> Nitrogen 14	8 <b>O</b> Oxygen 16	9 <b>F</b> Fluorine 19	10 <b>Ne</b> Neon 20
11 <b>Na</b> Sodium 23	12 <b>Mg</b> Magnesium 24											13 <b>Al</b> Aluminum 27	14 <b>Si</b> Silicon 28	15 <b>P</b> Phosphorus 31	16 <b>S</b> Sulphur 32	17 <b>Cl</b> Chlorine 35	18 <b>Ar</b> Argon 40
19 <b>K</b> Potassium 39	20 <b>Ca</b> Calcium 40	21 <b>Sc</b> Scandium 45	22 <b>Ti</b> Titanium 48	23 <b>V</b> Vanadium 51	24 <b>Cr</b> Chromium 52	25 <b>Mn</b> Manganese 55	26 <b>Fe</b> Iron 56	27 <b>Co</b> Cobalt 59	28 <b>Ni</b> Nickel 58	29 <b>Cu</b> Copper 63	30 <b>Zn</b> Zinc 64	31 <b>Ga</b> Gallium 69	32 <b>Ge</b> Germanium 74	33 <b>As</b> Arsenic 75	34 <b>Se</b> Selenium 80	35 <b>Br</b> Bromine 79	36 <b>Kr</b> Krypton 84
37 <b>Rb</b> Rubidium 85	38 <b>Sr</b> Strontium 88	39 <b>Y</b> Yttrium 89	40 <b>Zr</b> Zirconium 90	41 <b>Nb</b> Niobium 93	42 <b>Mo</b> Molybdenum 98	43 <b>Tc</b> Technetium 97	44 <b>Ru</b> Ruthenium 102	45 <b>Rh</b> Rhodium 103	46 <b>Pd</b> Palladium 106	47 <b>Ag</b> Silver 107	48 <b>Cd</b> Cadmium 114	49 <b>In</b> Indium 115	50 <b>Sn</b> Tin 120	51 <b>Sb</b> Antimony 121	52 <b>Te</b> Tellurium 130	53 <b>I</b> Iodine 127	54 <b>Xe</b> Xenon 132
55 <b>Cs</b> Caesium 133	56 <b>Ba</b> Barium 138	57-71	72 <b>Hf</b> Hafnium 180	73 <b>Ta</b> Tantalum 181	74 <b>W</b> Tungsten 184	75 <b>Re</b> Rhenium 187	76 <b>Os</b> Osmium 192	77 <b>Ir</b> Iridium 193	78 <b>Pt</b> Platinum 195	79 <b>Au</b> Gold 197	80 <b>Hg</b> Mercury 202	81 <b>Tl</b> Thallium 205	82 <b>Pb</b> Lead 208	83 <b>Bi</b> Bismuth 209	84 <b>Po</b> Polonium 209	85 <b>At</b> Astatine 210	86 <b>Rn</b> Radon 222
87 <b>Fr</b> Francium 223	88 <b>Ra</b> Radium 226	89-103	104 <b>Unq</b> Unnilquadium 260	105 <b>Unp</b> Unnilpentium 262	106 <b>Unh</b> Unnilhexium 263	107 <b>Uns</b> Unnilseptium 262	108 <b>Uno</b> Unniloctium 265	109 <b>Une</b> Unnilennium 266									
57 <b>La</b> Lanthanum 139	58 <b>Ce</b> Cerium 140	59 <b>Pr</b> Praseodymium 141	60 <b>Nd</b> Neodymium 142	61 <b>Pm</b> Promethium 145	62 <b>Sm</b> Samarium 152	63 <b>Eu</b> Europium 153	64 <b>Gd</b> Gadolinium 158	65 <b>Tb</b> Terbium 159	66 <b>Dy</b> Dysprosium 164	67 <b>Ho</b> Holmium 165	68 <b>Er</b> Erbium 168	69 <b>Tm</b> Thulium 169	70 <b>Yb</b> Ytterbium 174	71 <b>Lu</b> Lutetium 175			
89 <b>Ac</b> Actinium 227	90 <b>Th</b> Thorium 232	91 <b>Pa</b> Protactinium 231	92 <b>U</b> Uranium 238	93 <b>Np</b> Neptunium 237	94 <b>Pu</b> Plutonium 244	95 <b>Am</b> Americium 243	96 <b>Cm</b> Curium 247	97 <b>Bk</b> Berkelium 247	98 <b>Cf</b> Californium 251	99 <b>Es</b> Einsteinium 254	100 <b>Fm</b> Fermium 257	101 <b>Md</b> Mendelevium 258	102 <b>No</b> Nobelium 255	103 <b>Lr</b> Lawrencium 256			



# Why Neon? - (I) The Period Table of Trappable Elements

3 <b>Li</b> Lithium 7	4 <b>Be</b> Beryllium 9
11 <b>Na</b> Sodium 23	12 <b>Mg</b> Magnesium 24
19 <b>K</b> Potassium 39	20 <b>Ca</b> Calcium 40
37 <b>Rb</b> Rubidium 85	38 <b>Sr</b> Strontium 88
55 <b>Cs</b> Caesium 133	56 <b>Ba</b> Barium 138
87 <b>Fr</b> Francium 223	88 <b>Ra</b> Radium 226

24  
**Cr**  
Chromium  
52

47  
**Ag**  
Silver  
107

2  
**He**  
Helium  
4

10  
**Ne**  
Neon  
20

18  
**Ar**  
Argon  
40

36  
**Kr**  
Krypton  
84

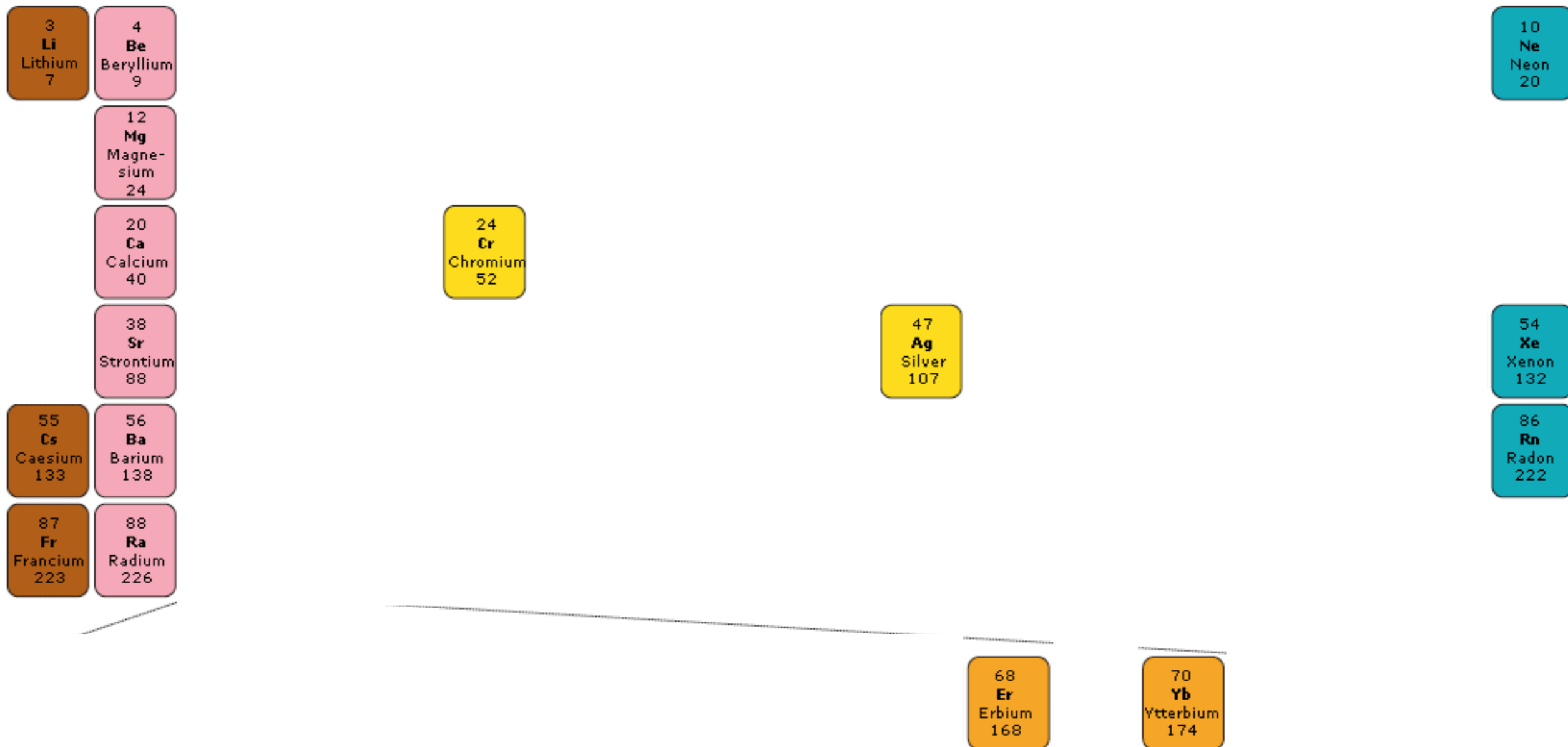
54  
**Xe**  
Xenon  
132

86  
**Rn**  
Radon  
222

68  
**Er**  
Erbium  
168

70  
**Yb**  
Ytterbium  
174

Why Neon? - (I) The Period Table of Trappable Elements Not already used for MOT  $\beta$  Decay



Why Neon? - (I) The Period Table of Trappable Elements Not already used for MOT  $\beta$  Decay & Decent to Calculate/Implement

3  
Li  
Lithium  
7

10  
Ne  
Neon  
20

Why Neon? - (I) The Period Table of Trappable Elements Not already used for MOT  $\beta$  Decay & Decent to Calculate/Implement

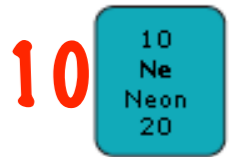


**Easier to produce**

**Trap setup exists in Israel (as does design for production of  $^8\text{Li}$ )**

**Easy wavelengths**

**Harder to trap (messy level scheme)**



**Easy to trap**

**Interesting physics other than beta decay**

**Needs to be in metastable state**

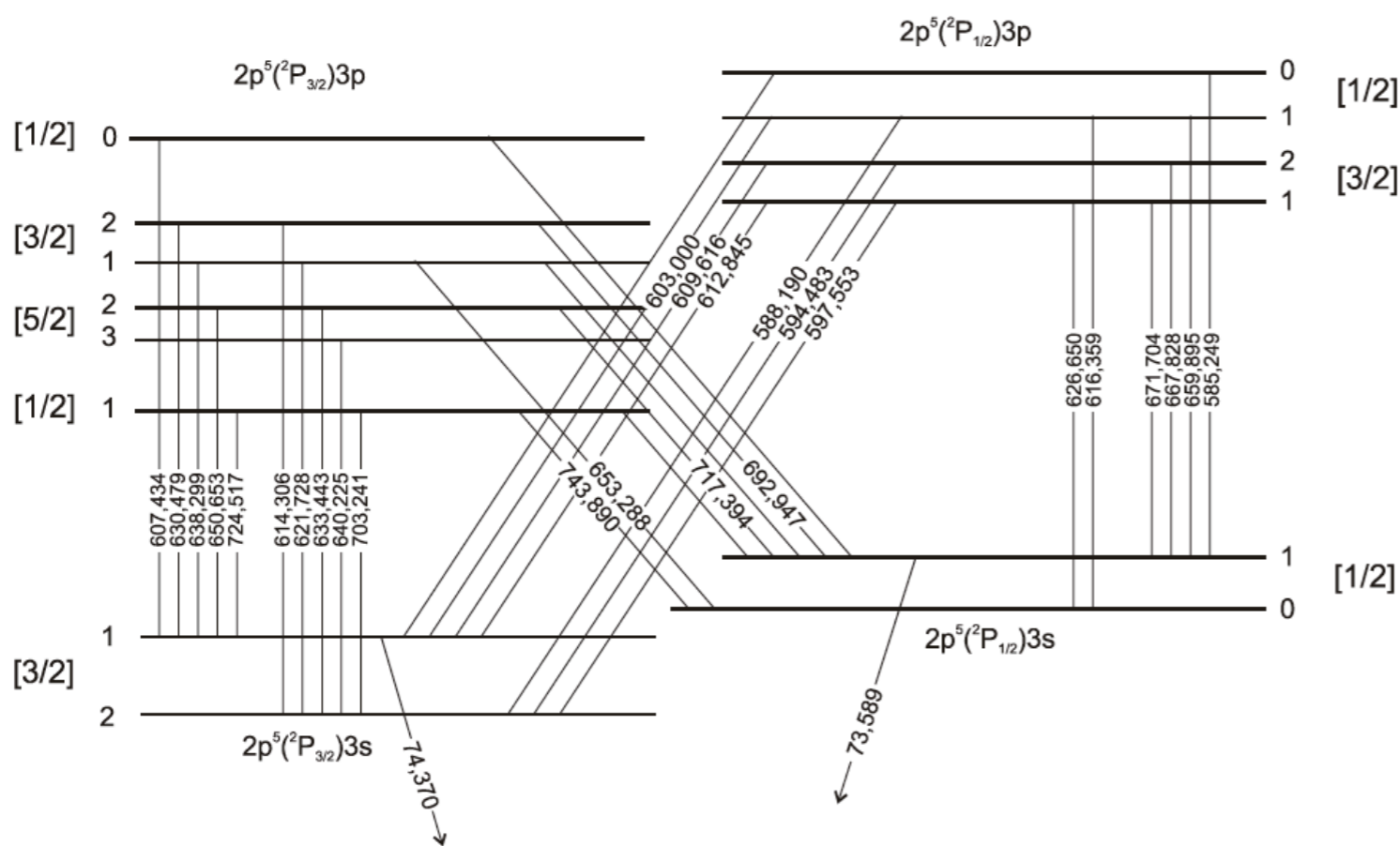
**Problematic WL for lasers**

**Harder to produce**

Why Neon? - (I) The Period Table of Trappable Elements Not already used for MOT  $\beta$  Decay & Decent to Calculate/Implement

10

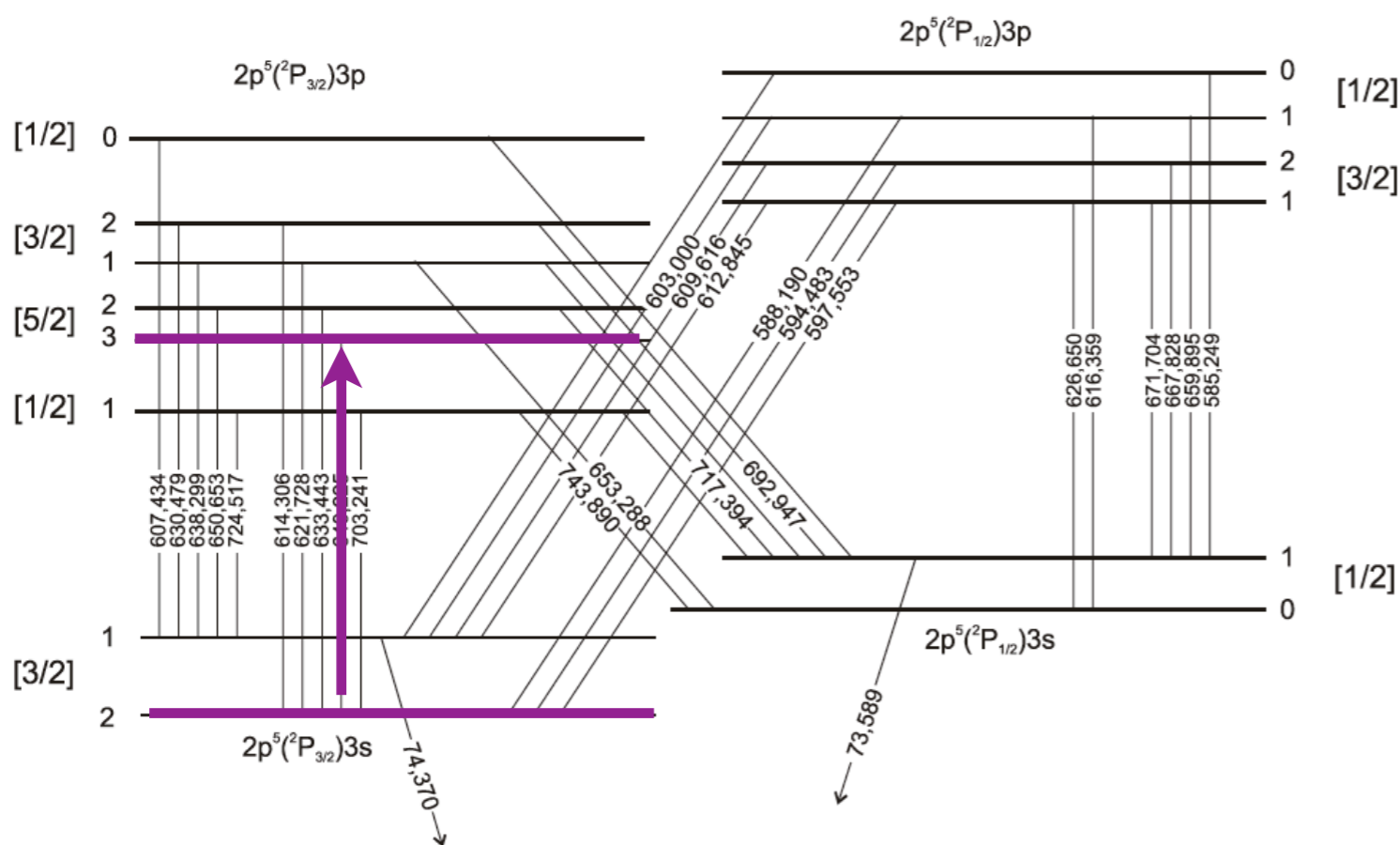
10  
Ne  
Neon  
20



Easy to trap  
 Interesting physics other the beta decay  
 Needs to be in metastable state  
 Problematic WL for lasers  
 Harder to produce

Why Neon? - (I) The Period Table of Trappable Elements Not already used for MOT  $\beta$  Decay & Decent to Calculate/Implement

10  
Ne  
Neon  
20



Fully closed transition for even neon isotopes (single frequency for trapping).

Easy to trap  
Interesting physics other the beta decay  
Needs to be in metastable state  
Problematic WL for lasers  
Harder to produce

Why Neon? - (I) The Period Table of Trappable Elements Not already used for MOT  $\beta$  Decay & Decent to Calculate/Implement

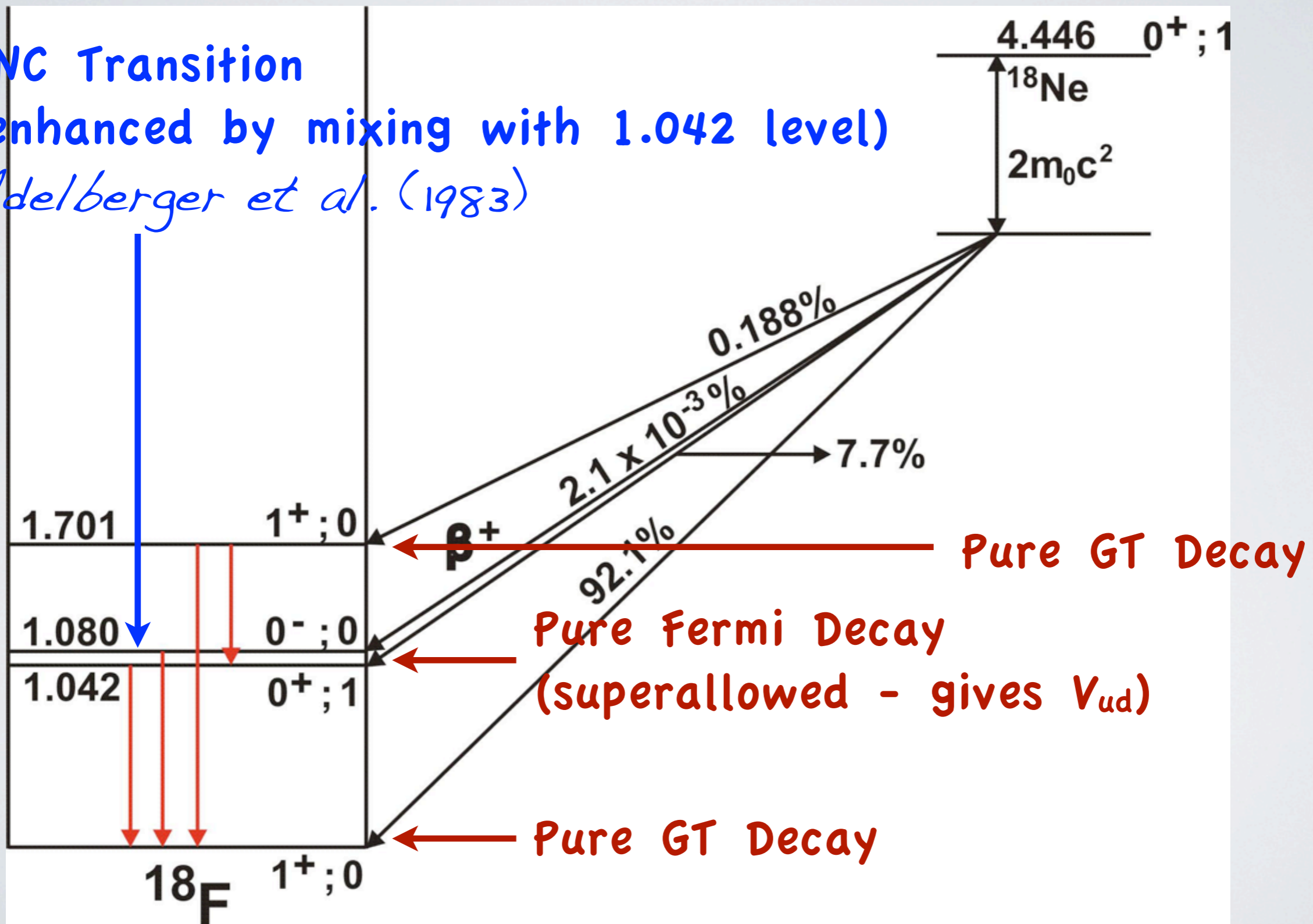
Ne -248.59° -246.08° -228.7° 0 20.1797 0.0112%	Ne16 122 keV 0+	Ne17 109.2 ms 1/2-	Ne18 1672 ms 0+	Ne19 17.22 s 1/2+	Ne20 0+	Ne21 3/2+	Ne22 0+	Ne23 37.24 s 5/2+	Ne24 3.38 m 0+	Ne25 602 ms (1/2,3/2)+	Ne26 197 ms 0+	Ne27 32 ms	Ne28 17 ms 0+	Ne29 0.2 s	Ne30 0+	Ne31	Ne32 0+
	2p	ECp,EC $\alpha$ ,...	EC	EC	90.48	0.27	9.25	$\beta^-$	$\beta^-$	$\beta^-$	$\beta n$	$\beta n$	$\beta n$	$\beta^-$			



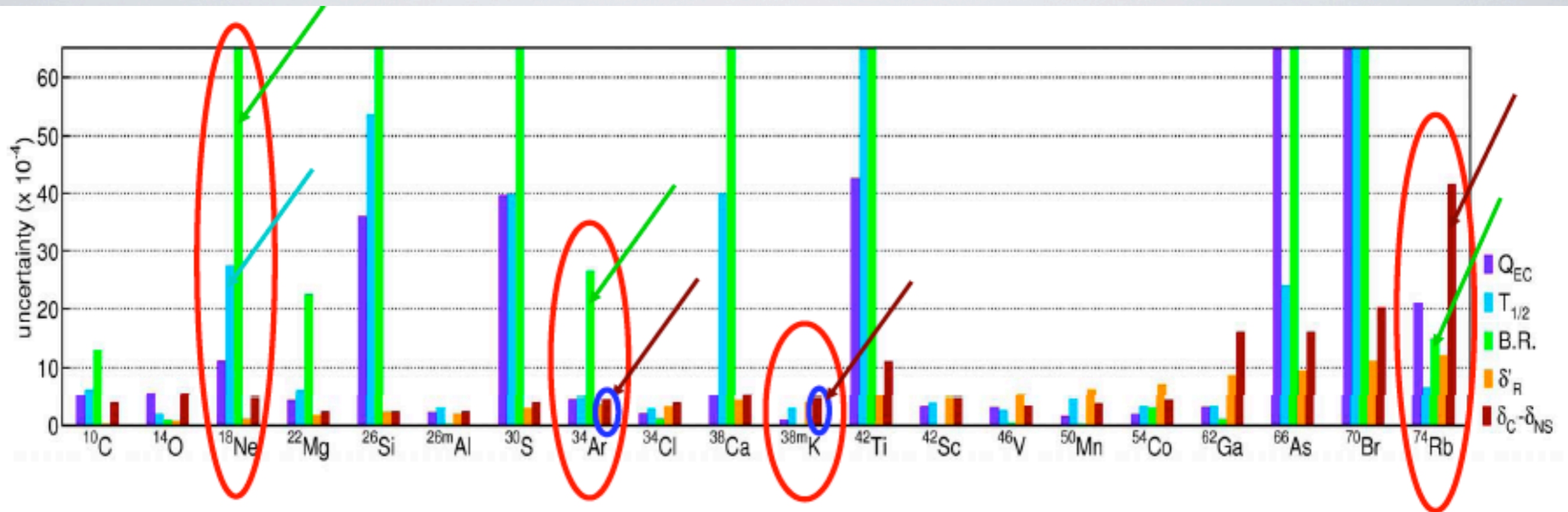


# Why $^{18}\text{Ne}$ ?

PNC Transition  
(enhanced by mixing with 1.042 level)  
*Adelberger et al. (1983)*



# $V_{ud}$ Uncertainties for superallowed decays



## Options

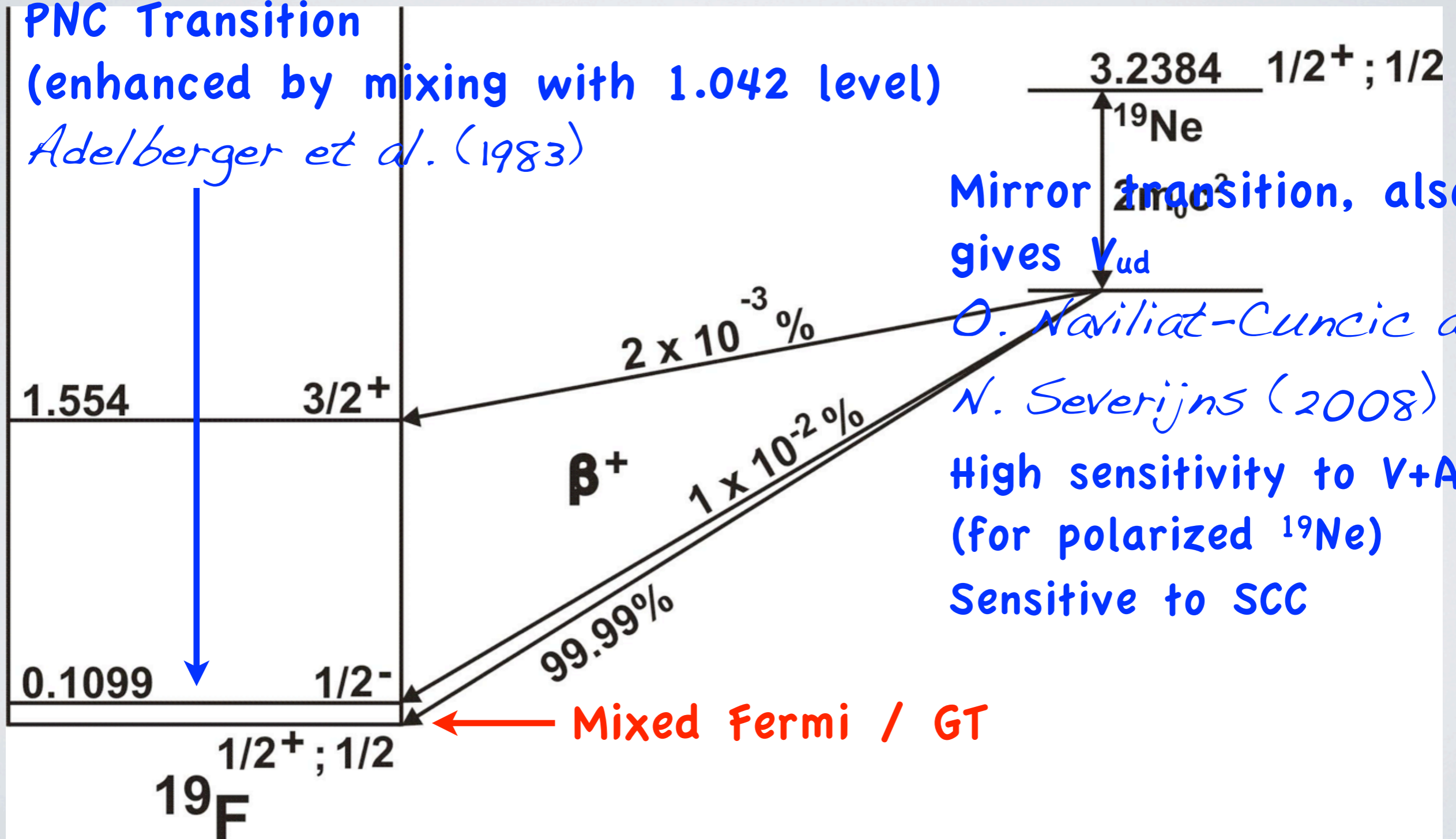
- improve quantities indicated by green & blue arrows
- if CVC accepted  $\rightarrow$  Ft-measurements test  $\delta_C - \delta_{NS}$  from theoretical models
- go for factor  $\sim 10$  higher precision in Ft than available now for the 4 isotopes indicated

# Why $^{19}\text{Ne}$ ?

## PNC Transition

(enhanced by mixing with 1.042 level)

*Adelberger et al. (1983)*



Mirror transition, also gives  $V_{ud}$

*O. Naviliat-Cuncic and N. Severijns (2008)*

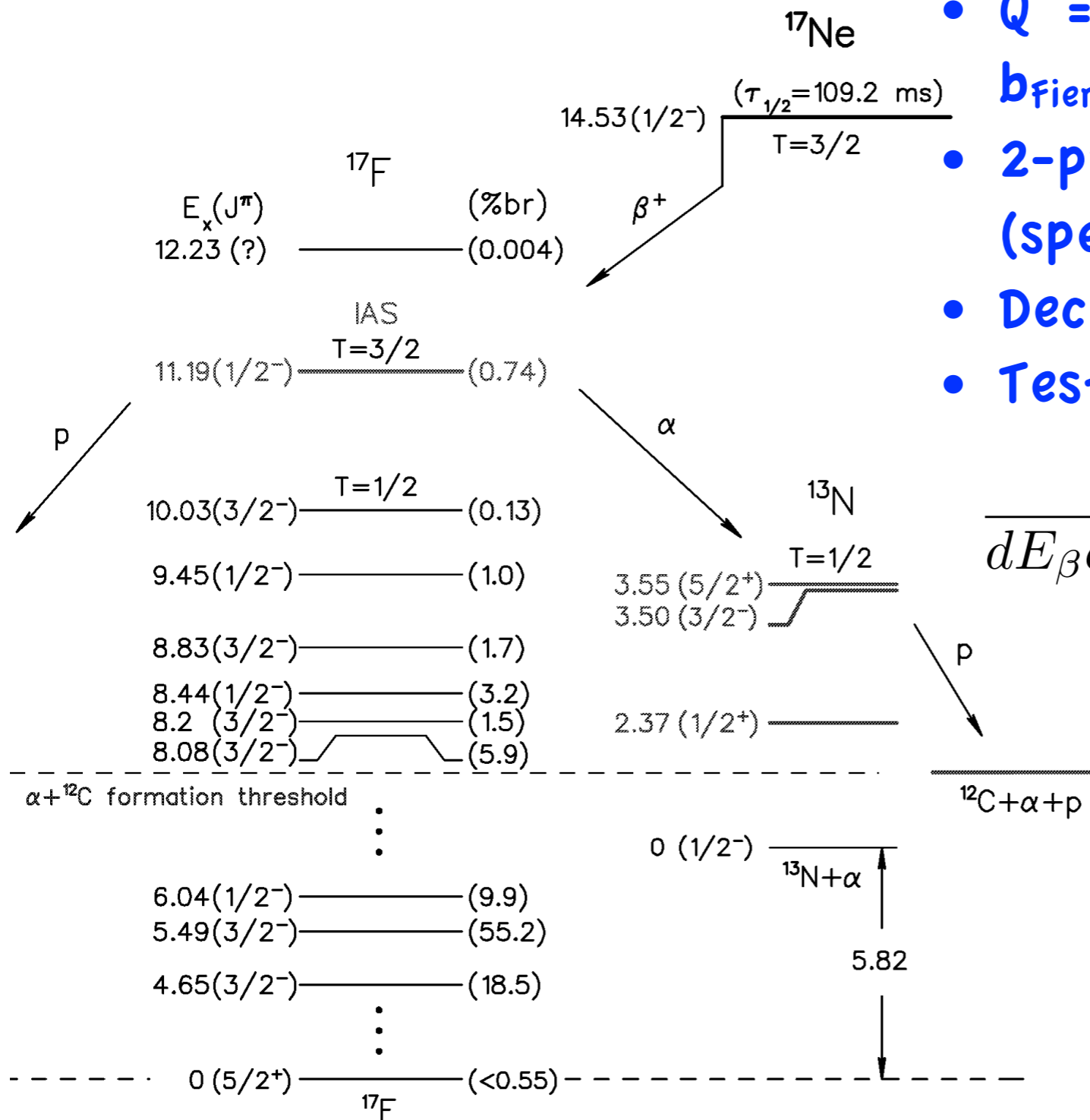
High sensitivity to  $V+A$  (for polarized  $^{19}\text{Ne}$ )

Sensitive to SCC

Mixed Fermi / GT

# Why $^{17}\text{Ne}$ ?

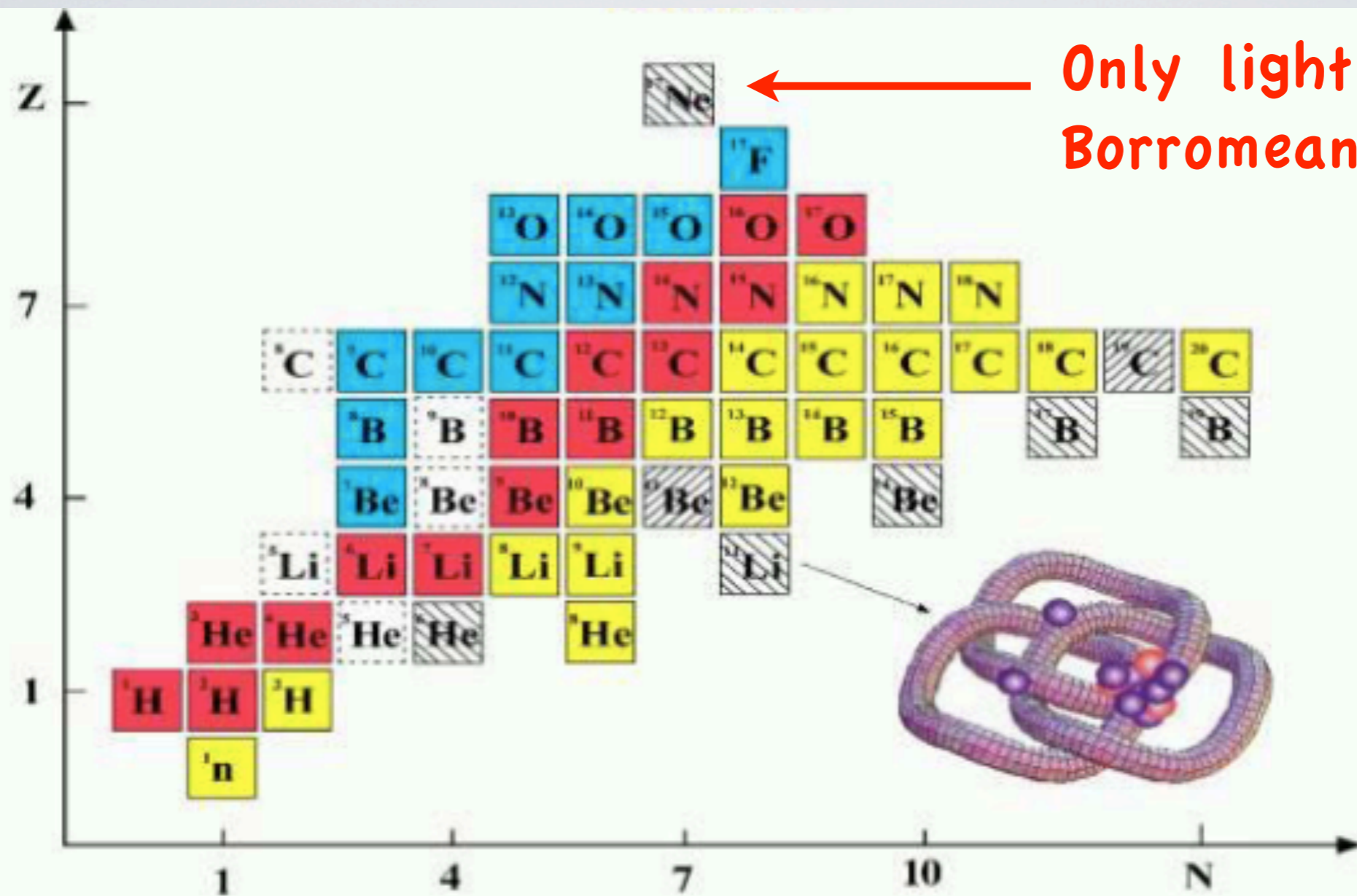
- $Q = 14\text{MeV}$  - allows extraction of  $b_{\text{Fierz}}$  (maybe)
- 2-proton Borromean halo nucleus (spectroscopy)
- Decay to halo nucleus  $^{17}\text{F}$
- Test of forbidden  $\beta$ -decays




$$\frac{d\Gamma}{dE_\beta d\Omega_\beta d\Omega_\nu} \propto \xi \left\{ 1 + a \frac{\vec{p}_e \cdot \vec{p}_\nu}{E_e E_\nu} + b \frac{m}{E_e} \right\}$$



**Borromean**



Only light proton  
Borromean nucleus

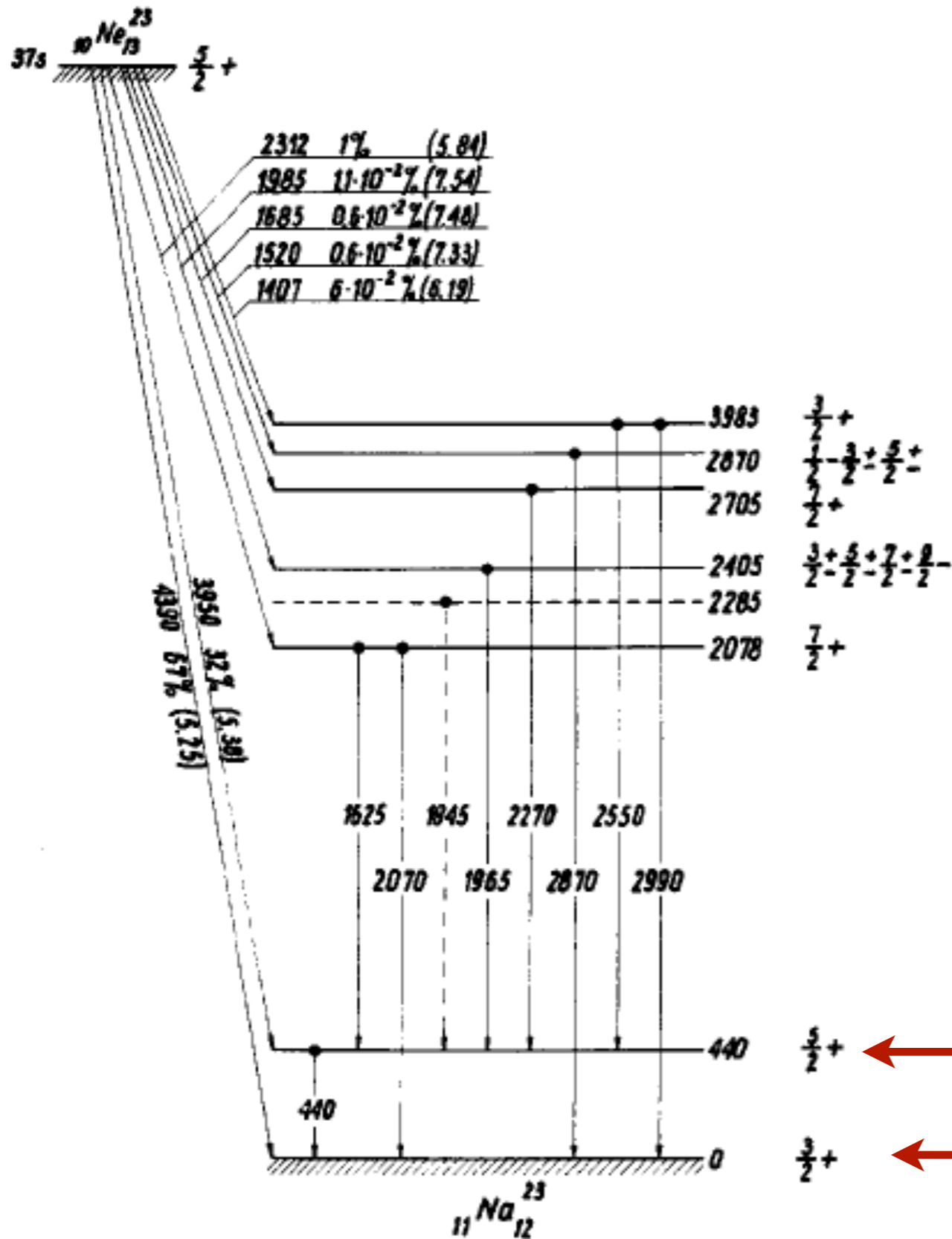
 Neutron halo

 Stable (~ 200)

 Borromean

 Unstable (> 6000)

# Why $^{23}\text{Ne}$ ?



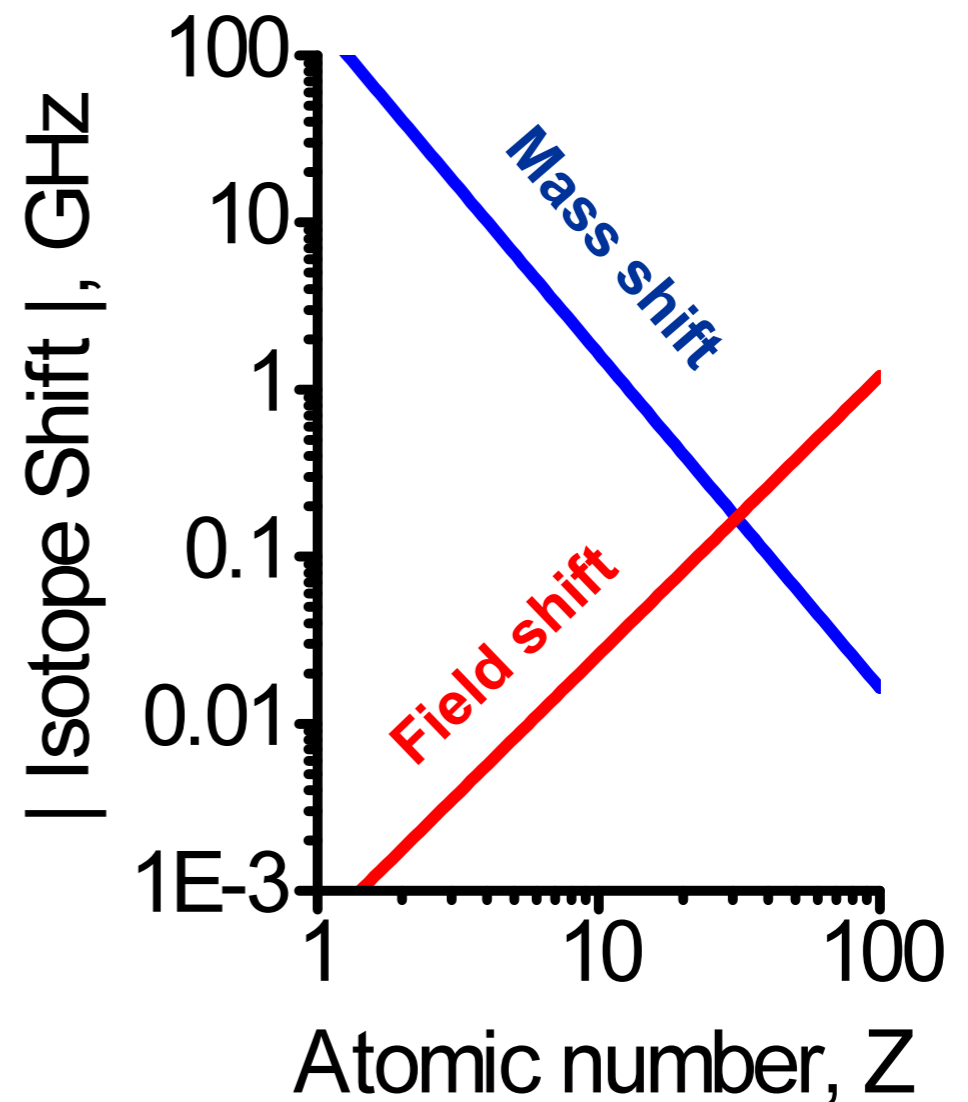
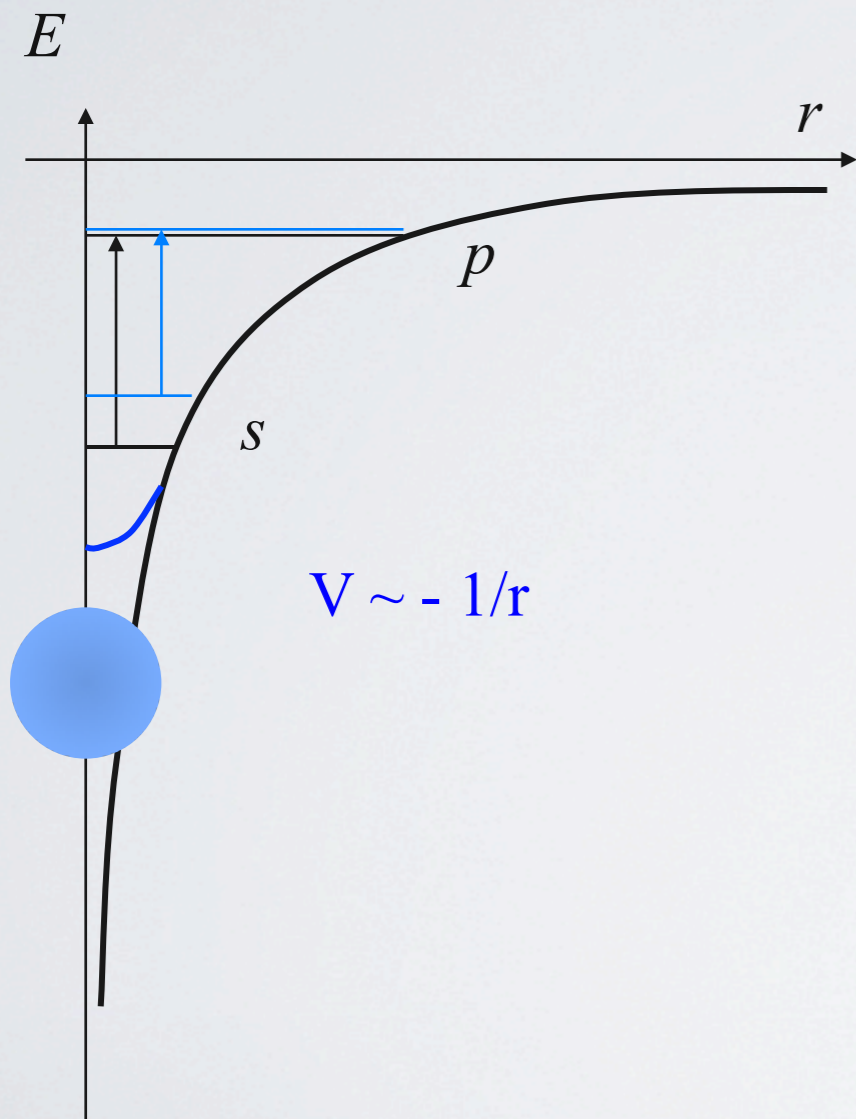
$$\frac{d\Gamma}{dE_\beta d\Omega_\beta d\Omega_\nu} \propto \xi \left\{ 1 + a \frac{\vec{p}_e \cdot \vec{p}_\nu}{E_e E_\nu} + b \frac{m}{E_e} \right\}$$

Almost Pure  
GT Decay  
Pure GT Decay

# NUCLEAR STRUCTURE TESTS WITH TRAPPED NEON ISOTOPES

## ISOTOPE SHIFTS

$$\delta\nu_{FS} = -\frac{2\pi}{3}Ze^2 \cdot \Delta |\psi(0)|^2 \cdot \delta \langle r^2 \rangle^{AA'}$$



Isotope Shift  
(Energy Level Change)

$$\delta\nu = \delta\nu_{FS} + \delta\nu_{MS}$$

Nuclear Size

$$\propto Z \cdot \delta \langle r^2 \rangle$$

CM Motion

$$\propto \frac{A - A'}{AA'}$$

## Neon vs. Helium *(Mueller et al.)* Mass Shift

Isotope	$(A' - A_{20}) / (A_{20} A')$
$^{17}\text{Ne}$	-0.0088
$^{18}\text{Ne}$	-0.0055
$^{19}\text{Ne}$	-0.0026
$^{21}\text{Ne}$	0.0023
$^{22}\text{Ne}$	0.0045
$^{23}\text{Ne}$	0.0065

Isotope	$(A' - A_{20}) / (A_{20} A')$
$^6\text{He}$	0.0833
$^8\text{He}$	0.125

- Sign change in Mass Shift.
- Effect  $\sim 10$  times smaller, better control on change in  $r^2$ .
- Harder to calculate for  $A > 12$ .
- But more cases (and 3 stable).



**It's not all fun and games**

# It's not all fun and games

- Neon trapping only in the excited metastable  $\text{Ne}^*$  state.
- RF discharge / electron beam excitation typically gives  $10^{-5}$  -  $10^{-6}$  efficiency.
- Neon trapping uses difficult wavelength (640nm).

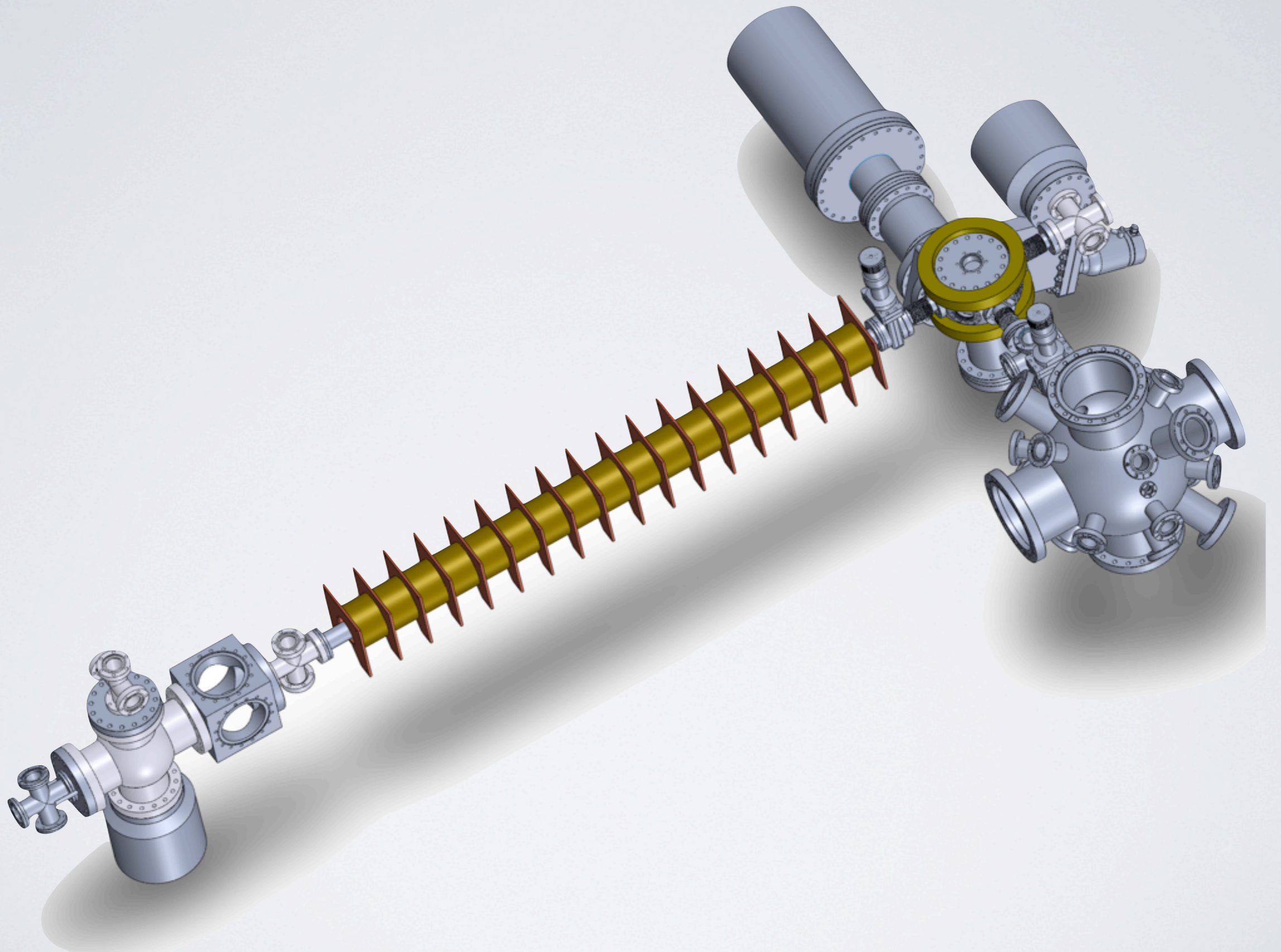
# It's not all fun and games

- Neon trapping only in the excited metastable  $\text{Ne}^*$  state.
- RF discharge / electron beam excitation typically gives  $10^{-5}$  -  $10^{-6}$  efficiency.
- Neon trapping uses difficult wavelength (640nm).

## But....

- Even numbers isotopes have no hyperfine splitting (single laser trapping) + Simple polarizing scheme for most odd isotopes ( $I=1/2$ ).
- Destructive probing of metastable beam is easy.
- Transport of noble gas from production target relatively simple (diffusion).
- Many isotopes to probe.
- Many combinations of coefficients  $\rightarrow$  many constraints on phase space.
- TOF calibration using Penning / Associative ionization  
$$\text{Ne}^* + \text{Ne}^* \rightarrow \text{Ne} + \text{Ne}^+ + e^-$$
$$\text{Ne}^* + \text{Ne}^* \rightarrow \text{Ne}_2^+ + e^-$$

# Some Technical Aspects



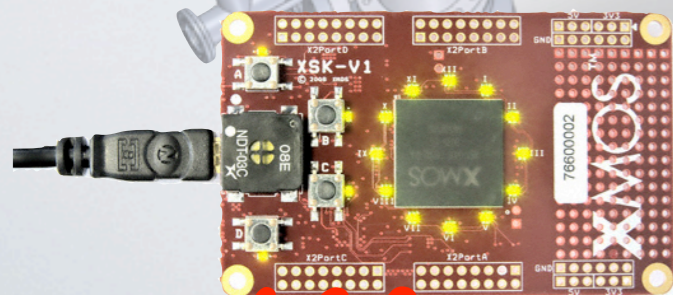
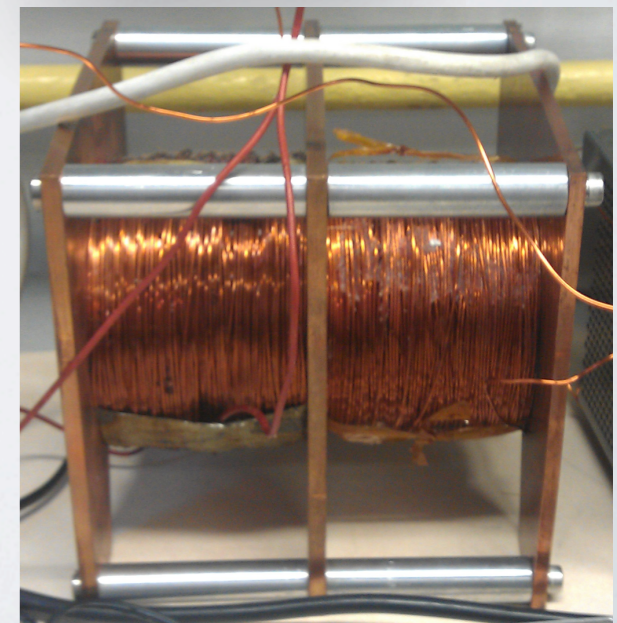
# Some Technical Aspects - I

## Multi-Coil Zeeman Slower

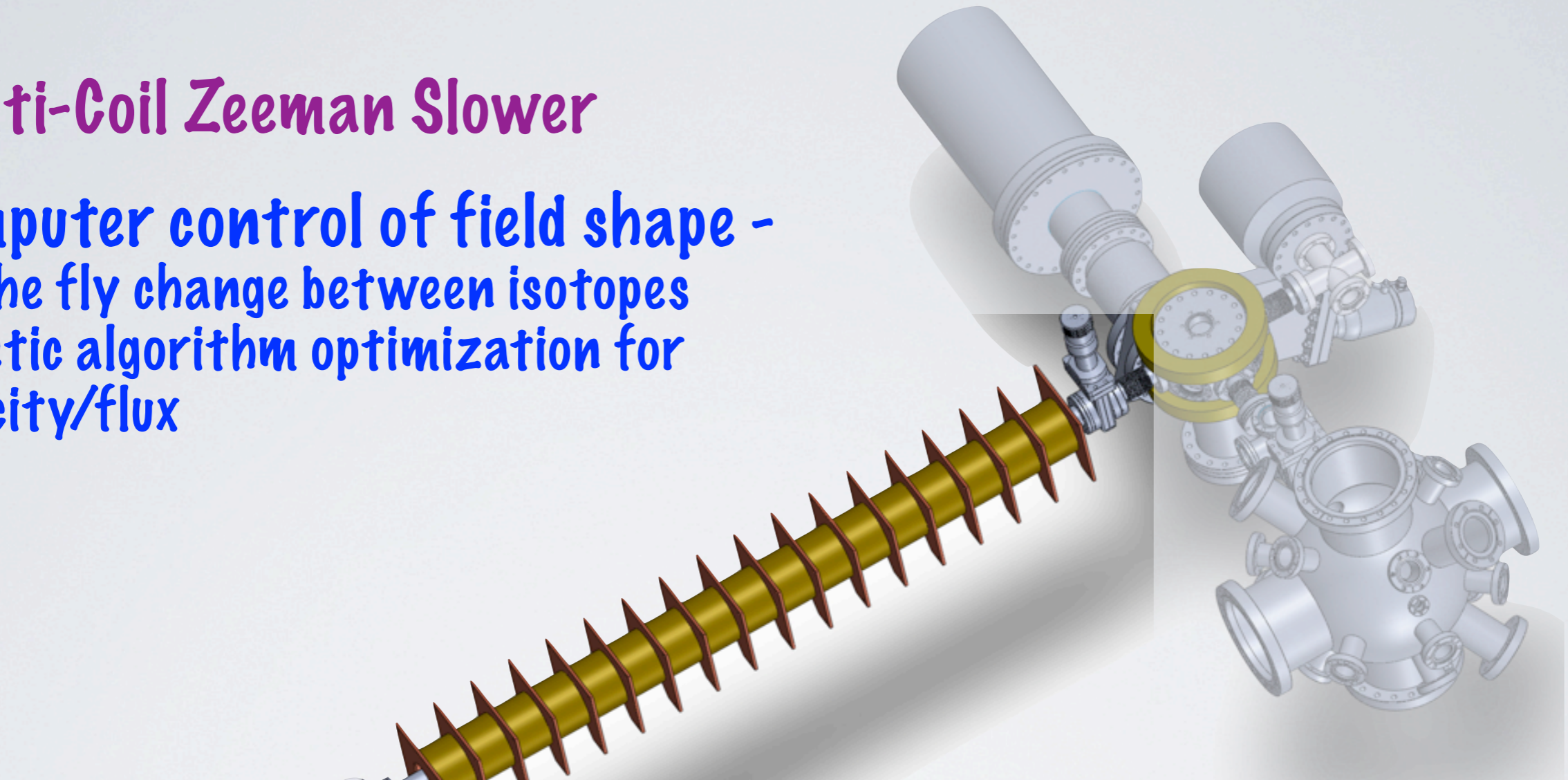
Computer control of field shape -  
On the fly change between isotopes  
Genetic algorithm optimization for  
velocity/flux

Field stability  $< 0.01\%$   
easily achieved

1/8 Coils



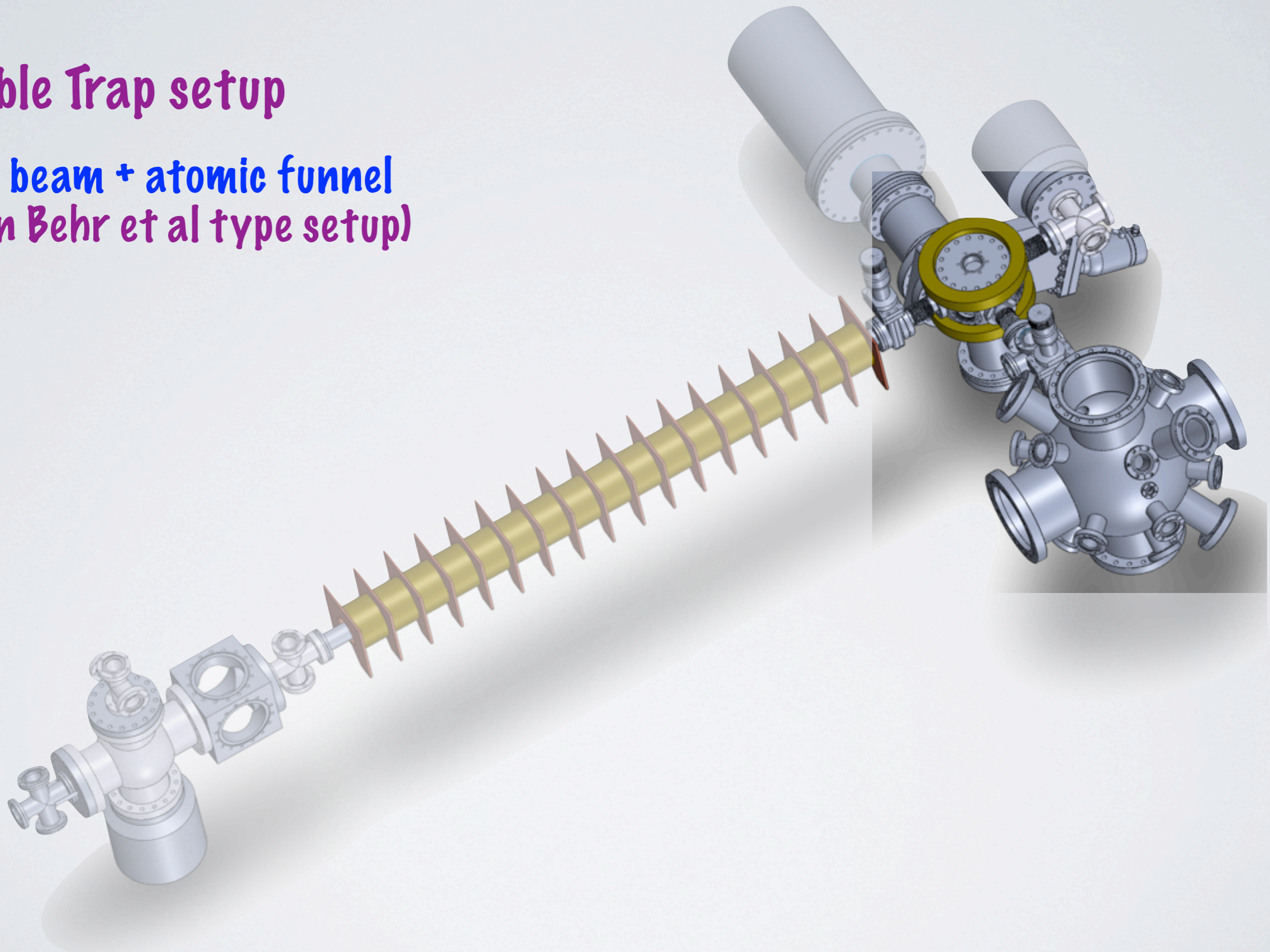
1-2 System



# Some Technical Aspects - II

## Double Trap setup

Push beam + atomic funnel  
(John Behr et al type setup)



# Some Technical Aspects - III

## “Faraday Cup” for Ne\* atoms

Ne\* atoms liberate electrons from metal plate -> Current flow

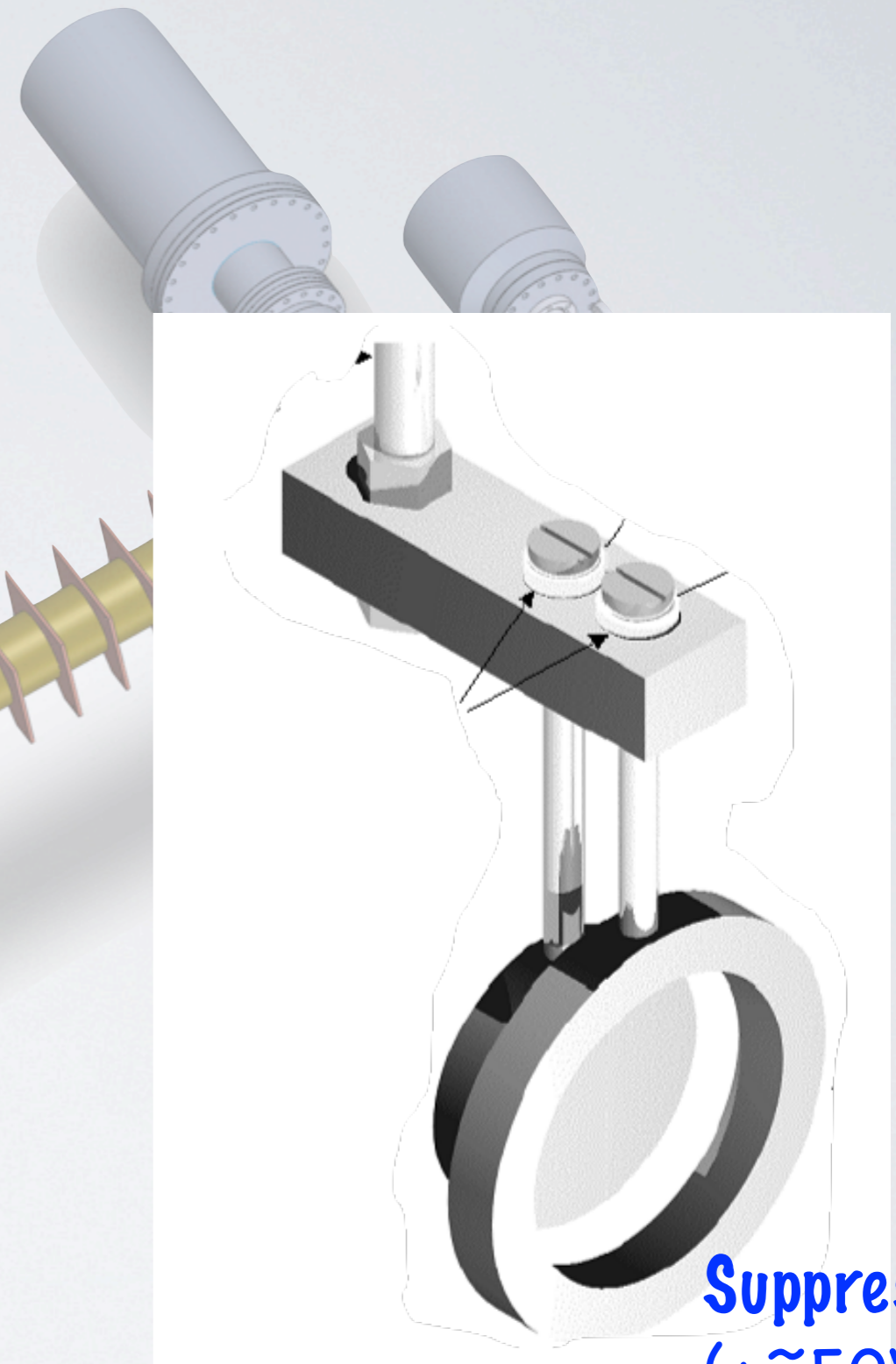
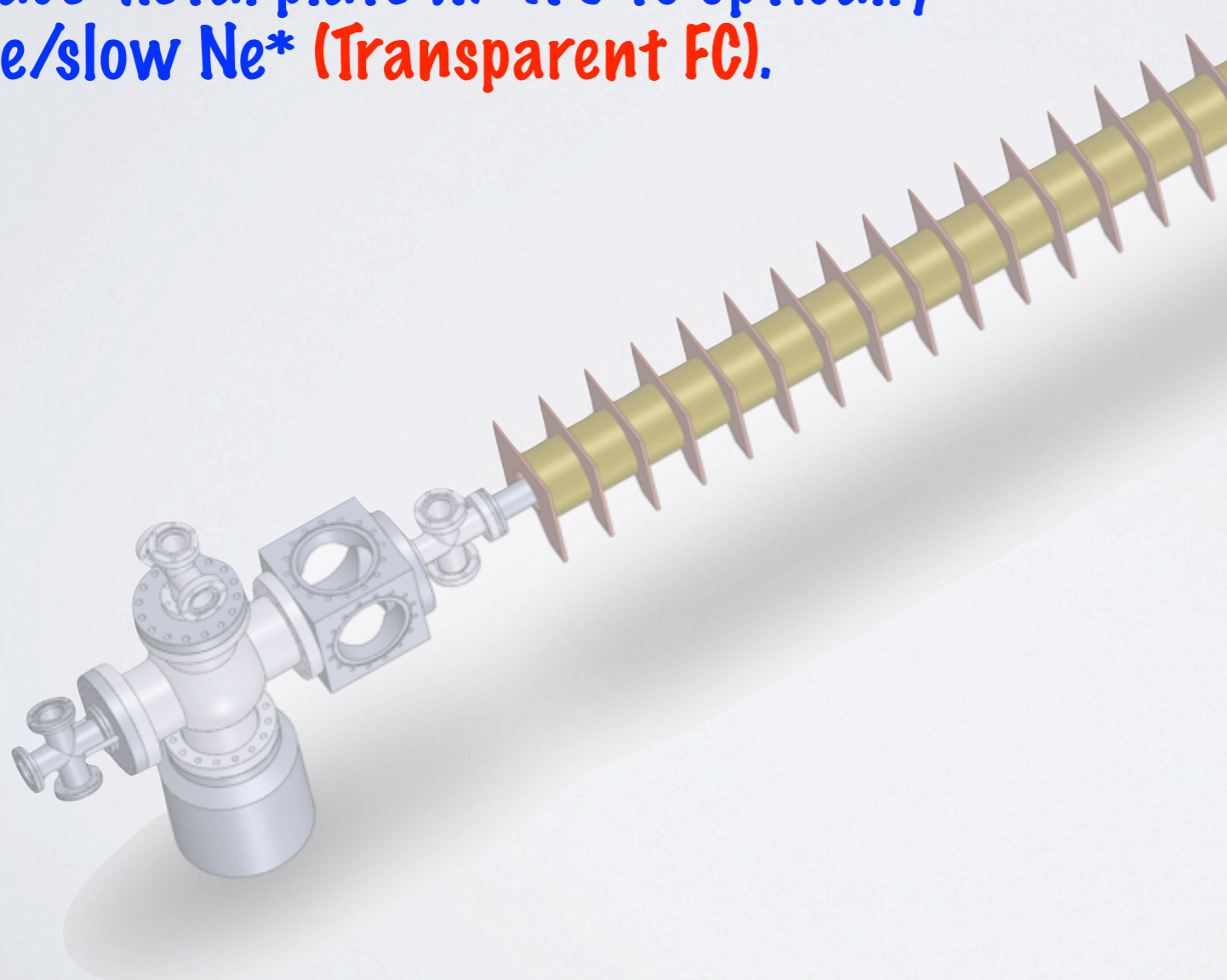


# Some Technical Aspects - III

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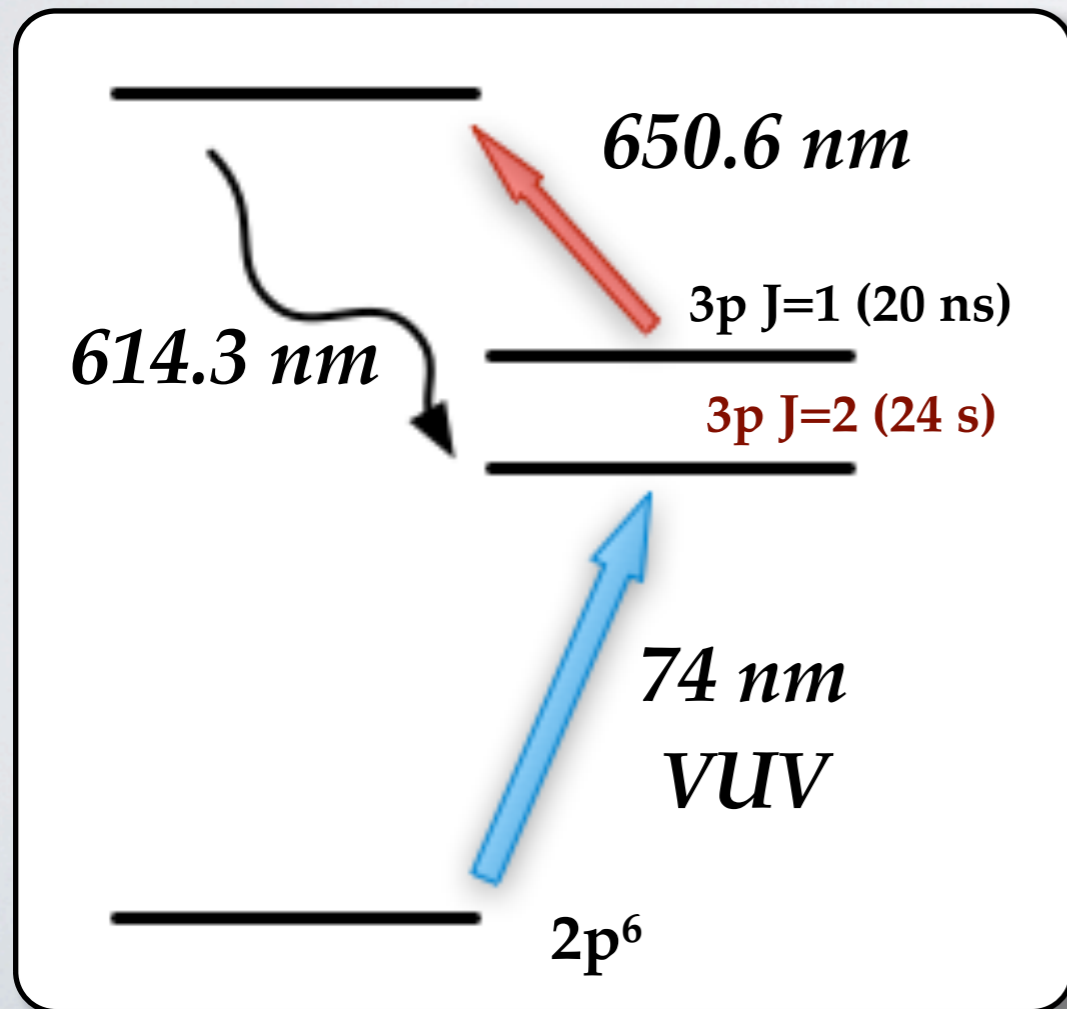
Replace metal plate w/ ITO to optically probe/slow Ne\* (Transparent FC).



Suppressor  
(+ ~50V)



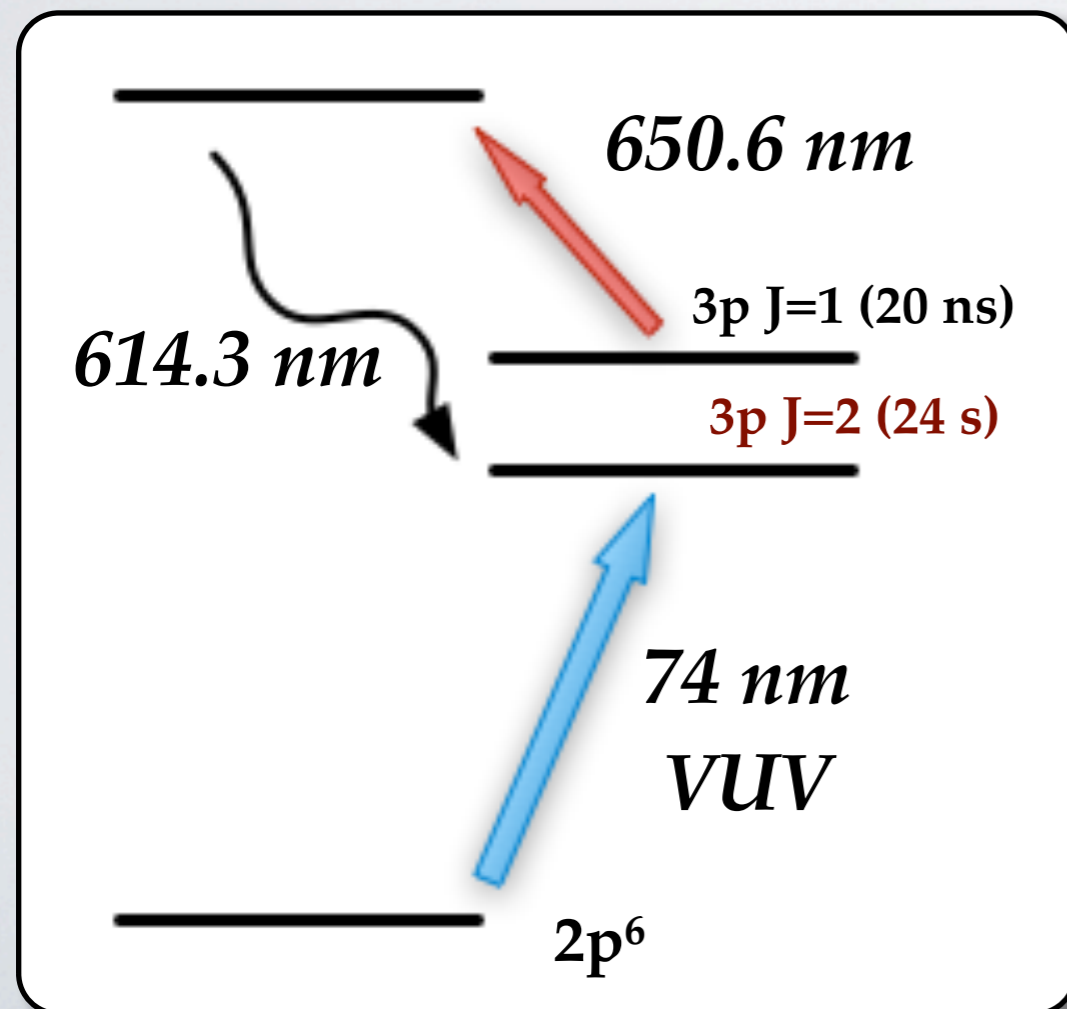
# Neon is Nice **But....** (Technical Aspect - IV)



# Neon is Nice **But....**

## (Technical Aspect - IV)

- Neon trapping only in the excited metastable  $\text{Ne}^*$  state.
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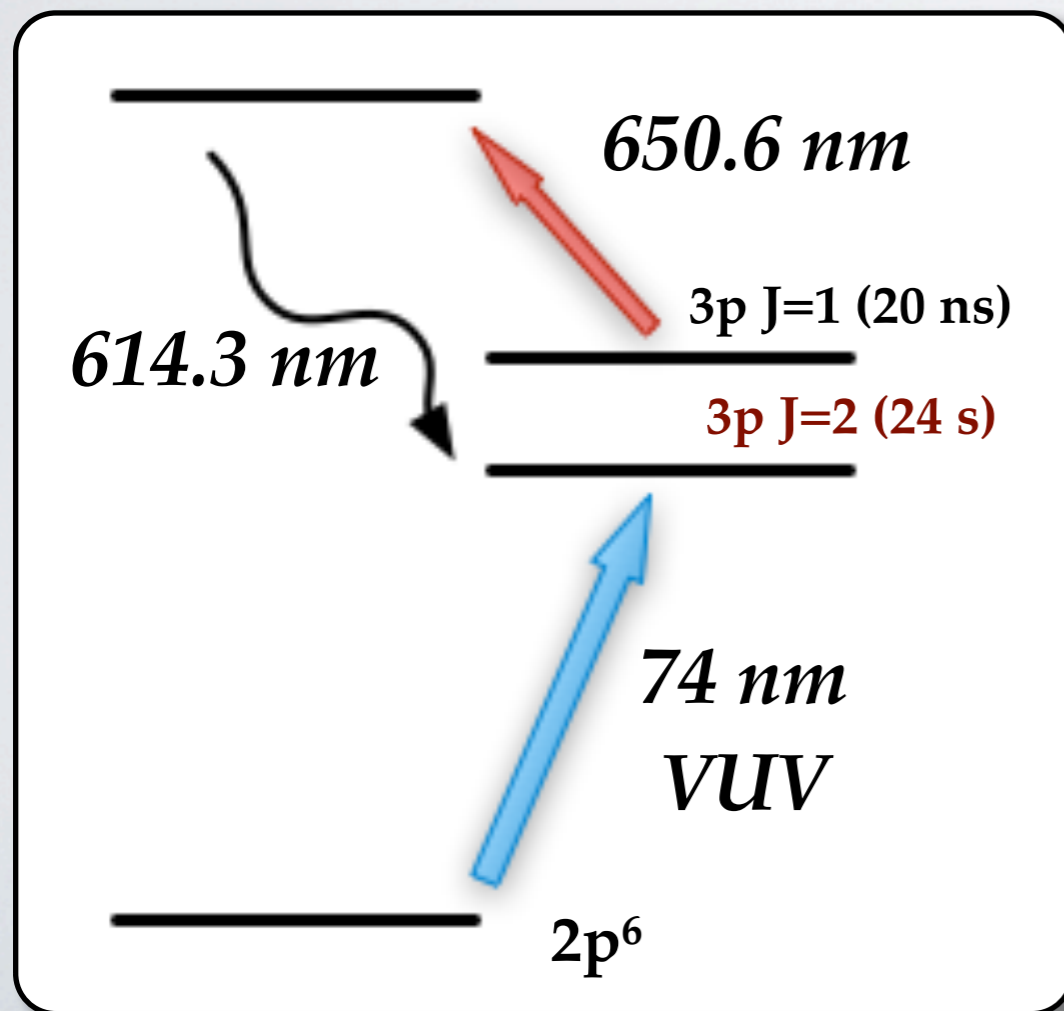
# Neon is Nice **But....**

## (Technical Aspect - IV)

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*Possible solution (based in work at ANL for Kr/Ar)*

*All optical excitation*



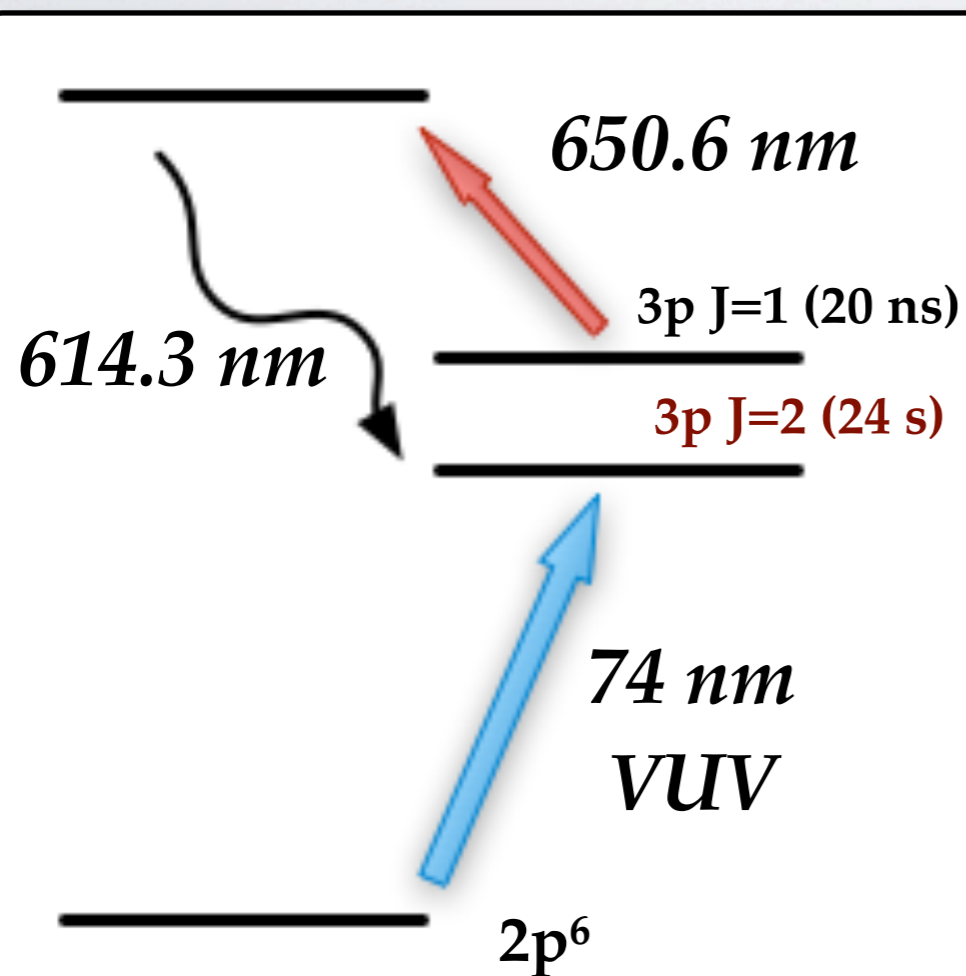
- 2 photon excitation and spontaneous decay into metastable (long lived) state.
- $^{18}\text{Ne}$  has same level structure as stable  $^{20}\text{Ne}$ , some complications with  $^{19}\text{Ne}$ .
- Trapping on  $J=2 \rightarrow J'=3$  transition at  $640 \text{ nm}$ .

# Neon is Nice **But....** (Technical Aspect - IV)

Predicted improved of 2-3 orders of magnitude in efficiency.

*Possible solution (based in work at ANL for Kr/Ar)*

*All optical excitation*



- 2 photon excitation and spontaneous decay into metastable (long lived) state.
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- Trapping on  $J=2 \rightarrow J'=3$  transition at  $640\text{nm}$ .

# Measurement Scheme

## General Scheme

Produce Ne Isotopes

Move to source chamber (TMP?)

Excite to  $\text{Ne}^*$  +  
expand into Slower

Take data  
(TOF, position  
asymmetry on MCP)

Move to science  
chamber (push  
beam)

Trap in collection  
trap

Continuous  
Intermittent



# A Brief Aside

## Optical traps

- Once cooled and trapped by the MOT, atoms can be trapped by the purely dipole force.

*Interaction of laser E field and induced dipole moment:*

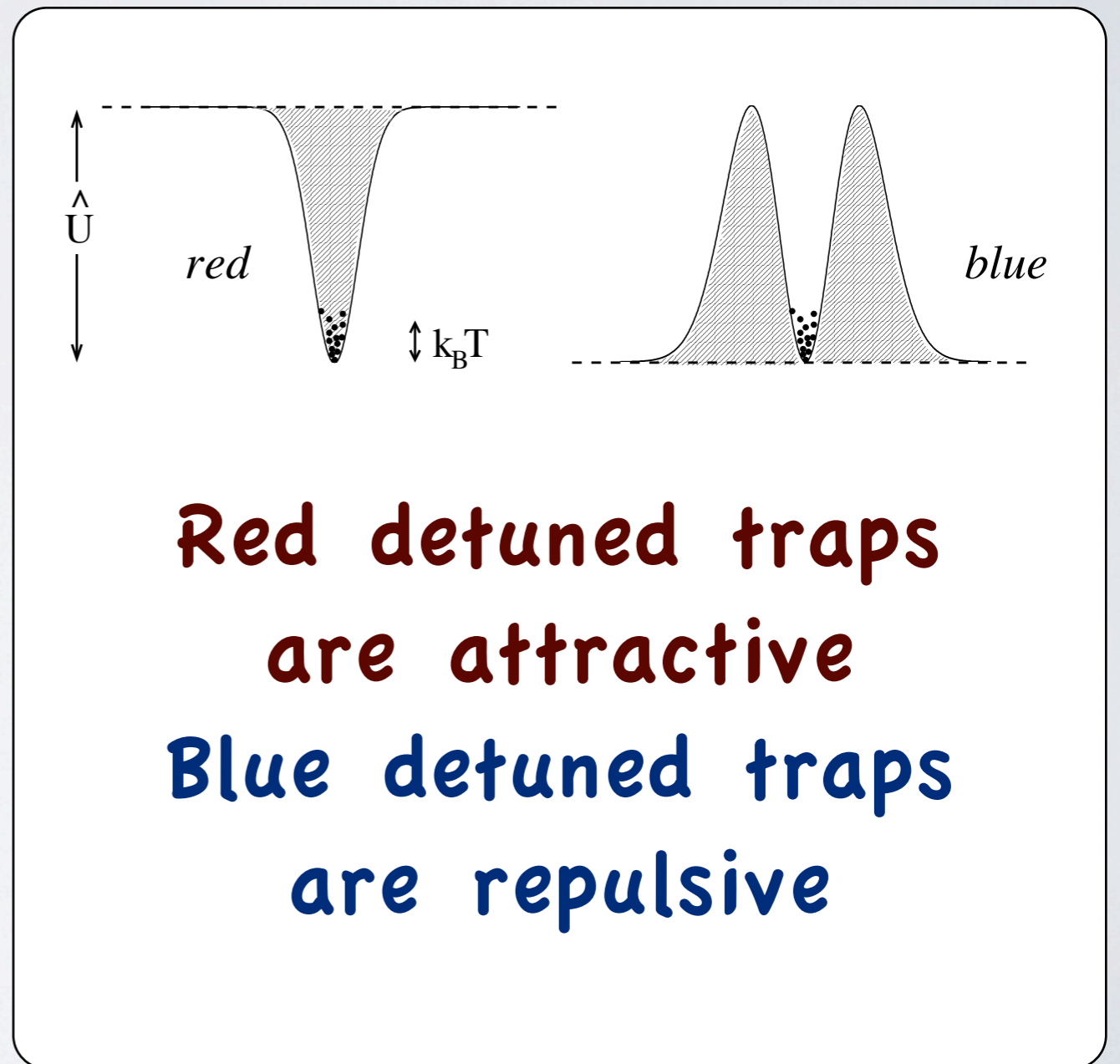
$$\tilde{\mathbf{p}} = \alpha \tilde{\mathbf{E}}$$

$$U_{dip} = -\frac{1}{2} \langle \mathbf{p} \mathbf{E} \rangle$$

$$U_{dip}(\vec{r}) = \frac{3\pi c^2}{2\omega_0^3} \frac{\Gamma}{\Delta} I(\vec{r})$$

$$\gamma_{sc}(\vec{r}) = \frac{3\pi c^2}{2\hbar\omega_0^3} \left( \frac{\Gamma}{\Delta} \right)^2 I(\vec{r})$$

$$\Delta = \omega_{laser} - \omega_0$$



# Looking Ahead (Tech V)

## Dark Blue Traps

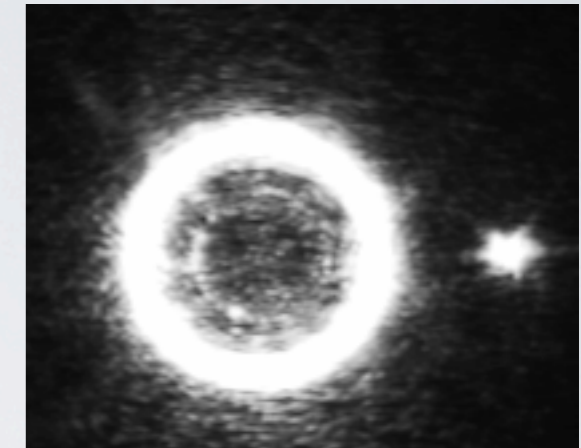
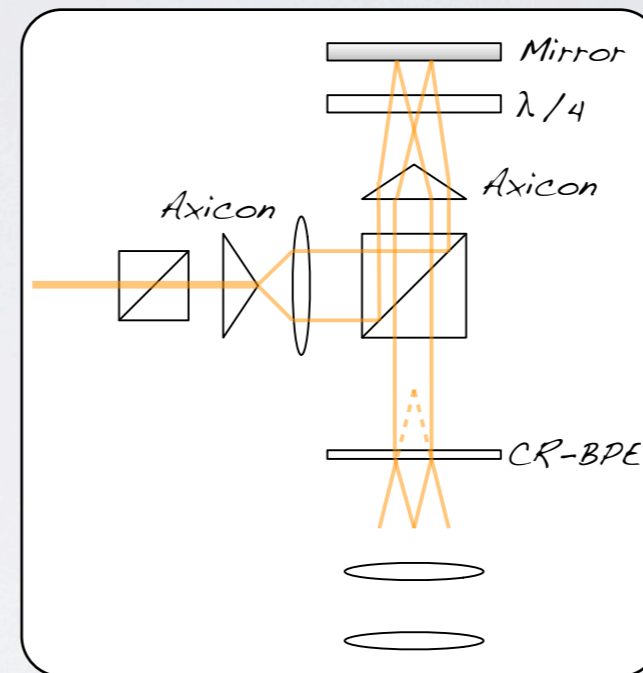
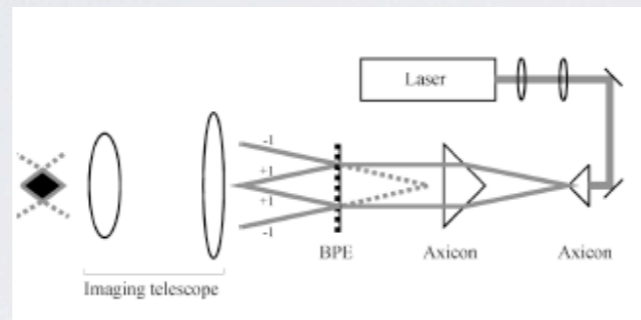
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Single beam  
“axicon” trap



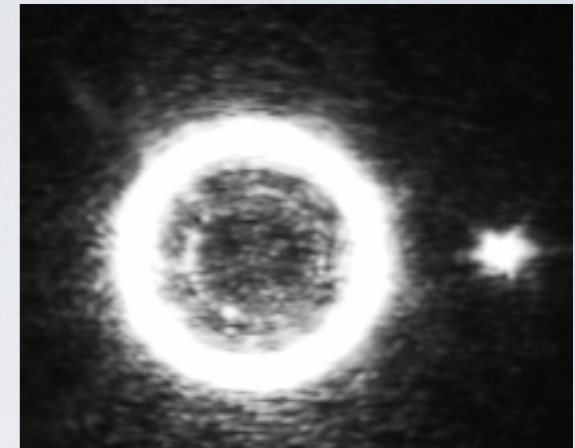
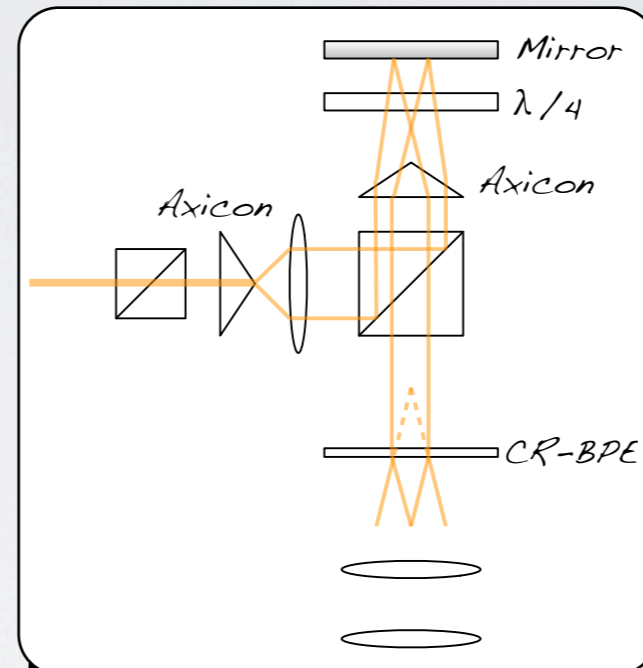
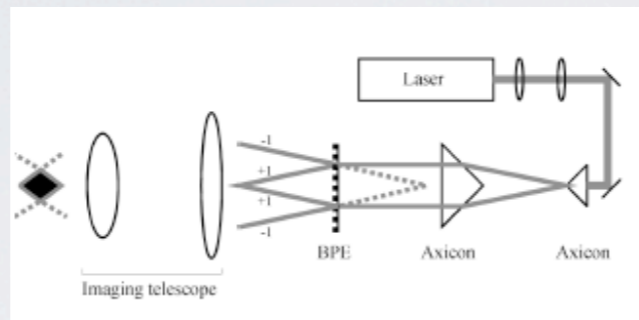


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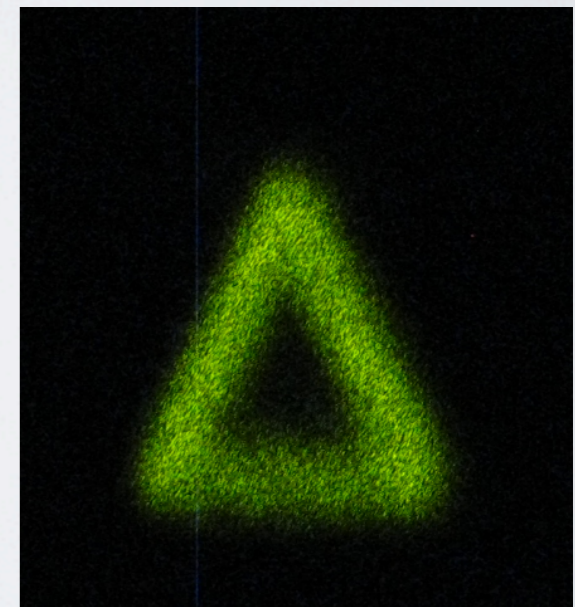
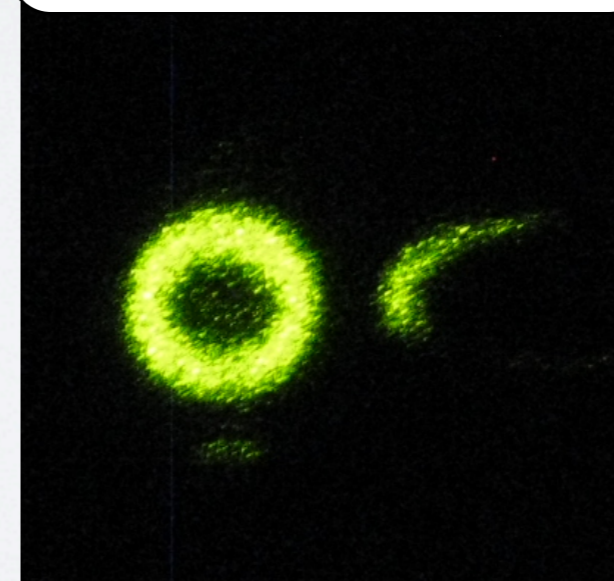
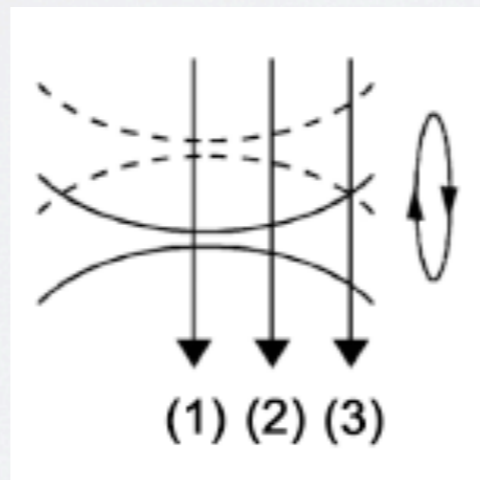
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ROtating Beam  
Optical Trap  
(but in 2 orthogonal  
directions)



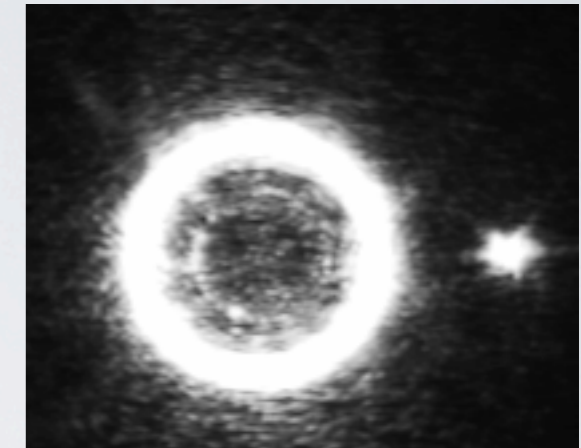
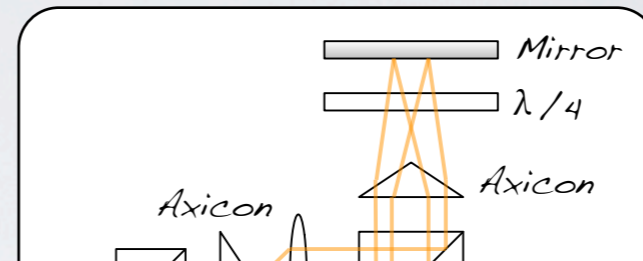
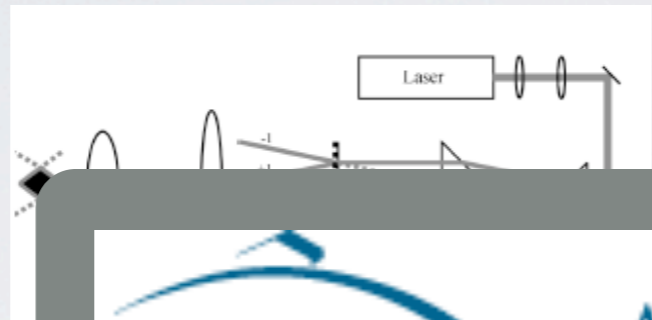
*Digitally controlled*

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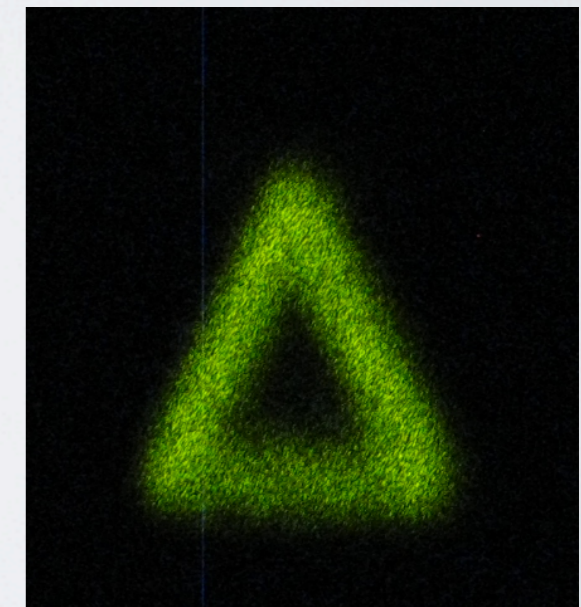
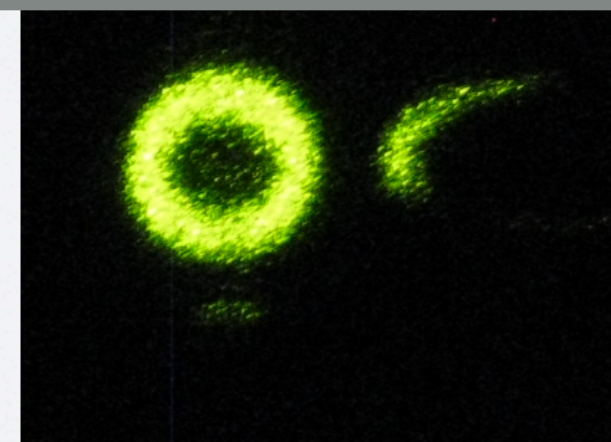
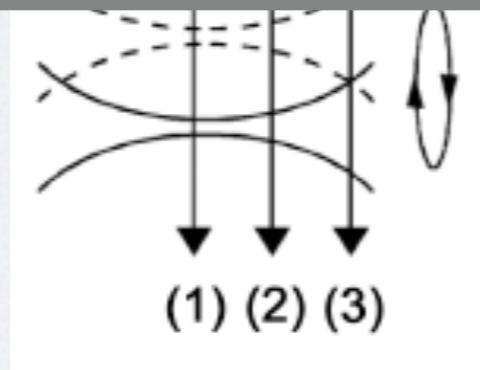
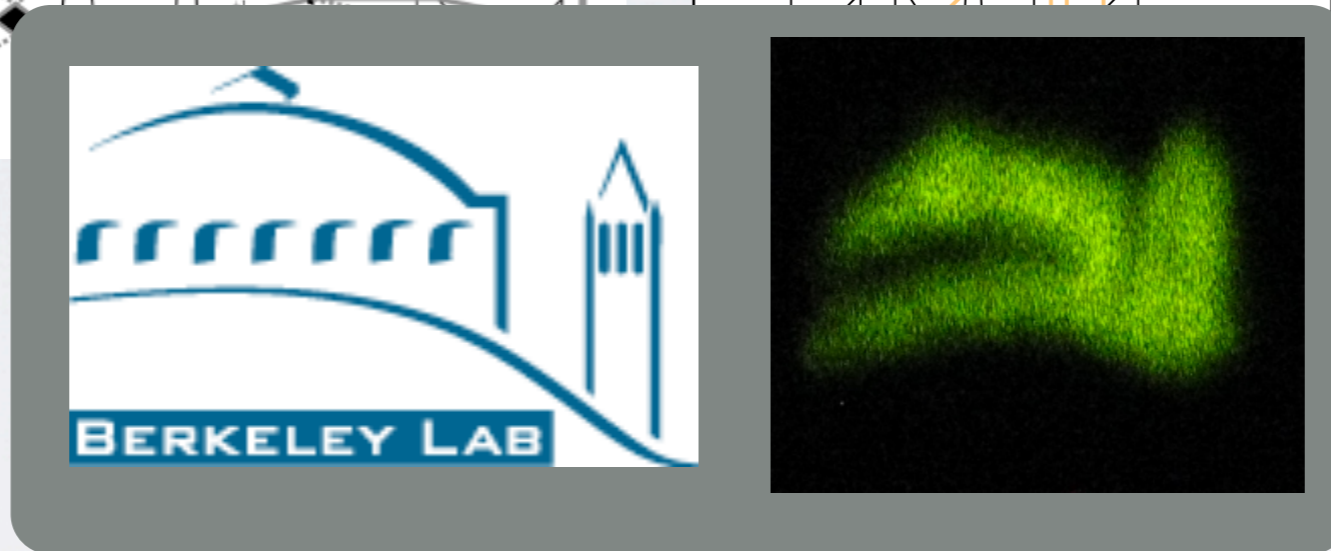
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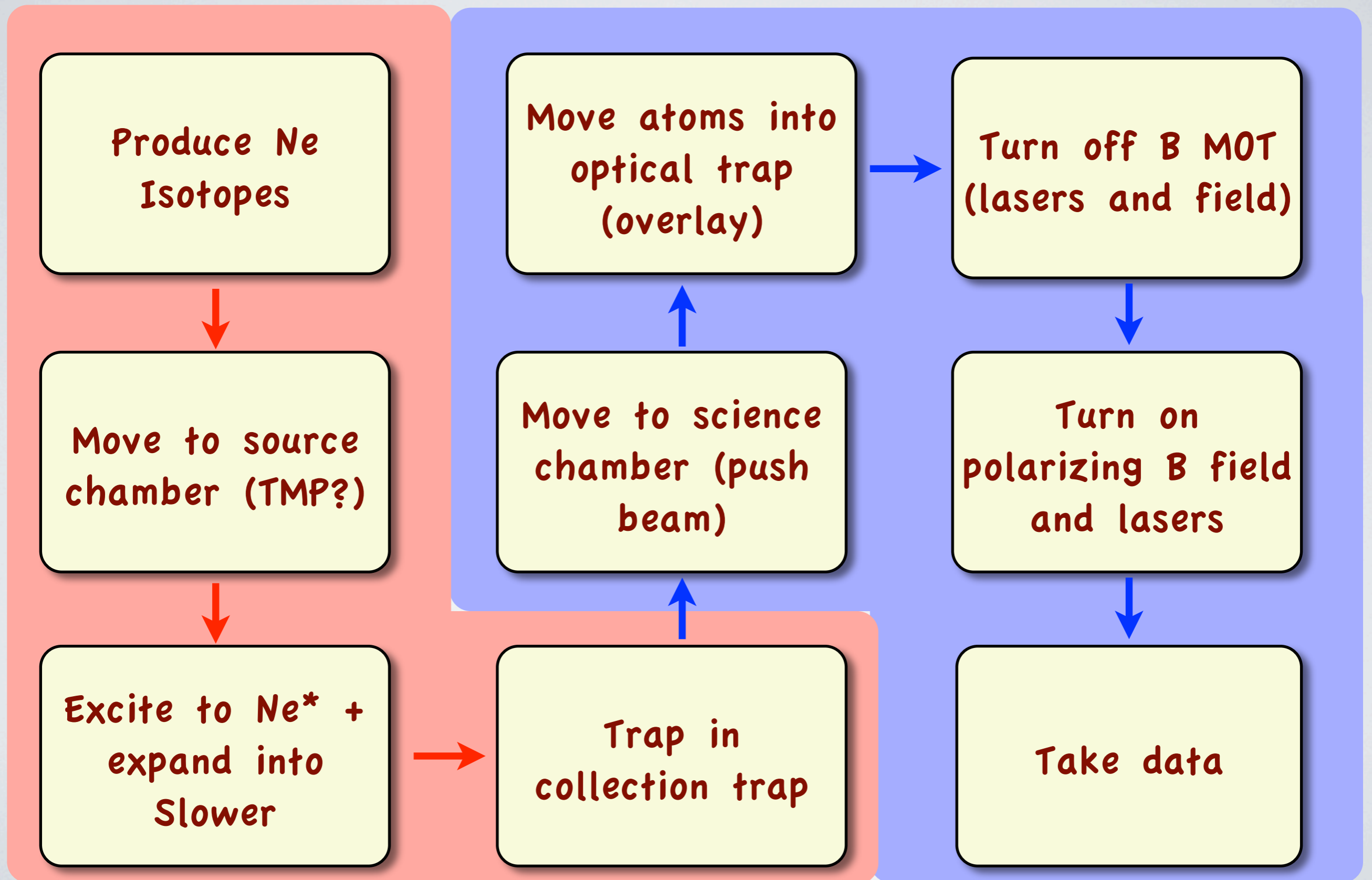
ROtating Beam  
Optical Trap  
(but in 2 orthogonal  
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*Digitally controlled*

# Polarization Measurement

## General Scheme

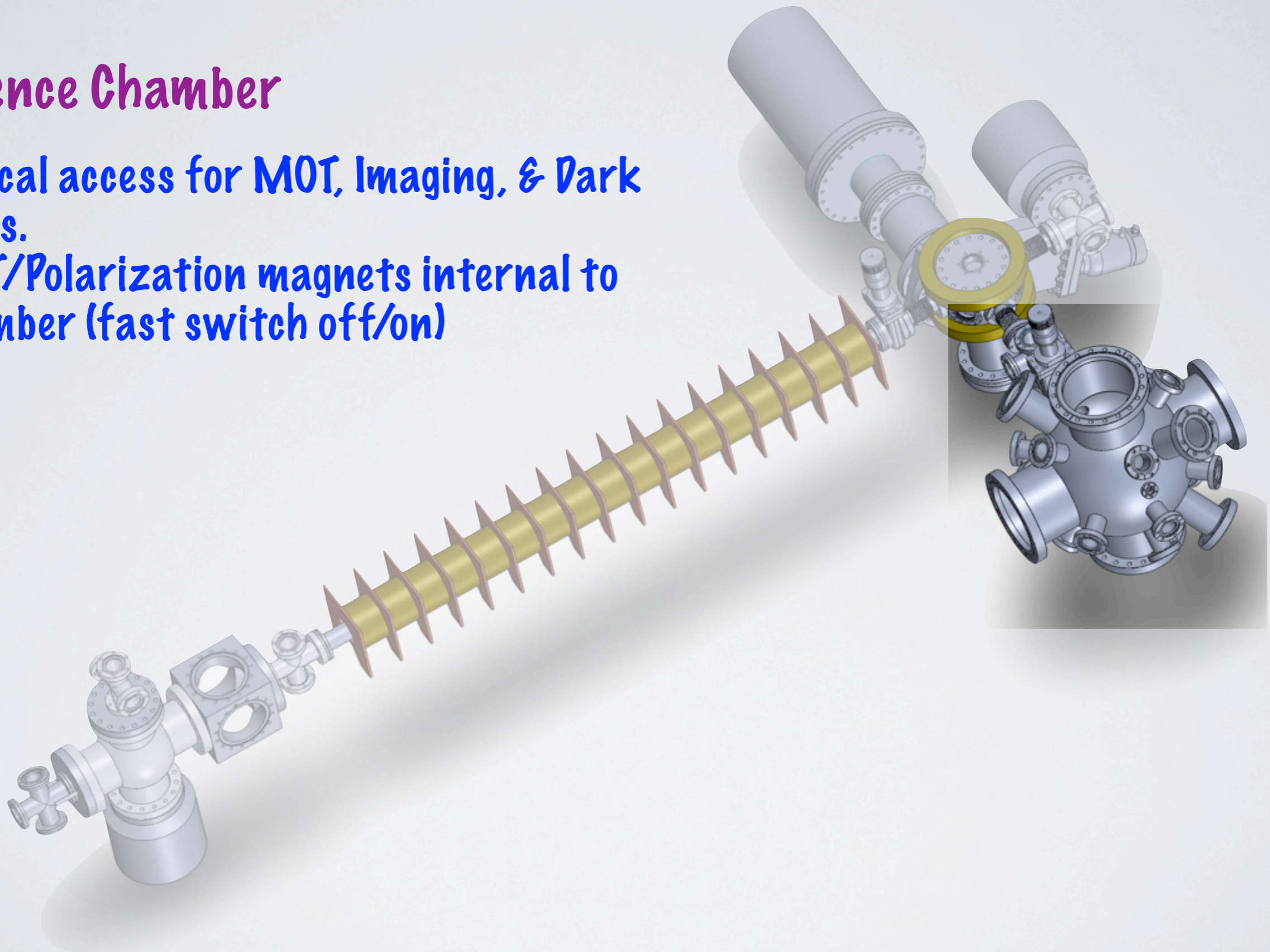


# Some Technical Aspects - VI

## Science Chamber

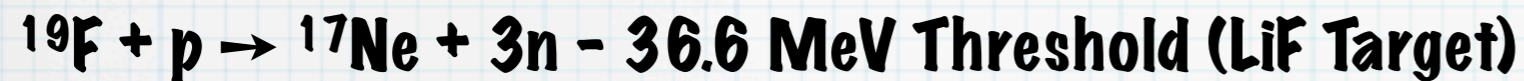
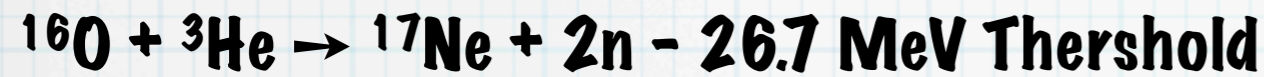
Optical access for MOT, Imaging, & Dark traps.

MOT/Polarization magnets internal to chamber (fast switch off/on)

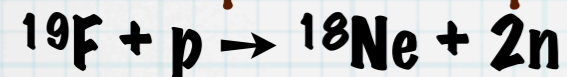


# PROSPECTS FOR PRODUCTION IN ISRAEL

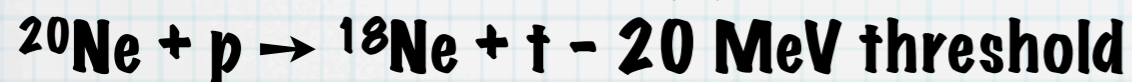
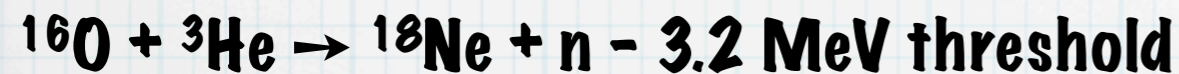
## $^{17}\text{Ne}$ :



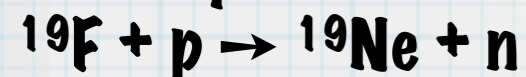
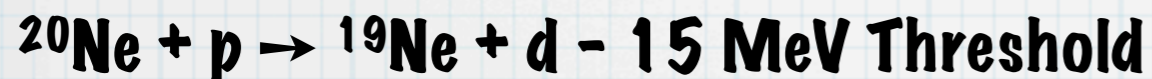
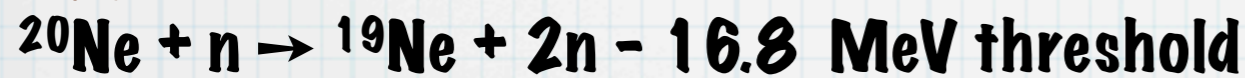
## $^{18}\text{Ne}$ : Explored as a possible source for $\beta$ beams



- 13 MeV Protons of LiF target,  $3 \times 10^{-5}/\text{p}$  (Takayama (2009))

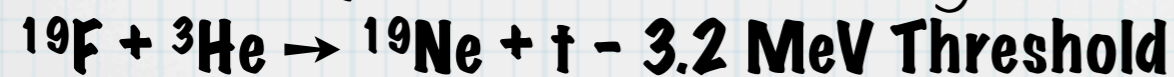


## $^{19}\text{Ne}$ :



- 4 MeV Threshold

- ( $10^{-3}/\text{p}$  for thick LiF, Kitwanga et. al (1990)).



## $^{23}\text{Ne}$ :



# CAN THIS BE DONE IN ISRAEL?

## SARAF

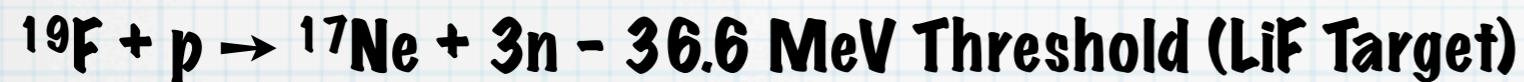
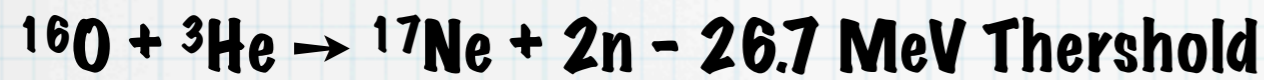
- \* New, (very) high current p/d accelerator (5mA/up to 40MeV) under construction at SOREQ.
- \* Neutron production also possible with Liquid-Li (for eg., but under construction). **Talk by M. Hass.**

## Weizmann Institute

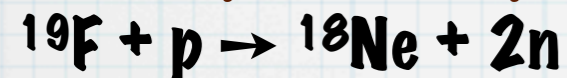
- \* Neutron generator being purchased.
- \* Some production schemes may also be possible with the VDG accelerator.

# PROSPECTS FOR PRODUCTION IN ISRAEL

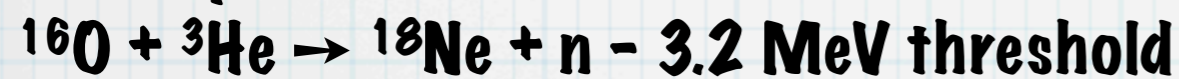
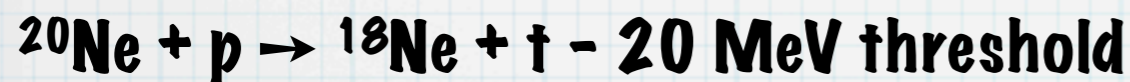
## **<sup>17</sup>Ne:**



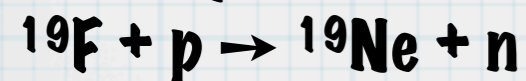
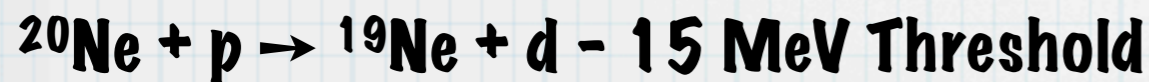
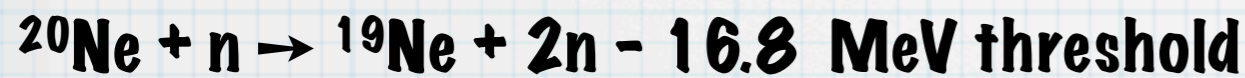
## **<sup>18</sup>Ne: Explored as a possible source for $\beta$ beams**



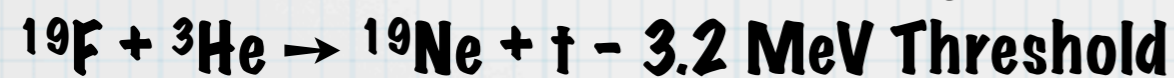
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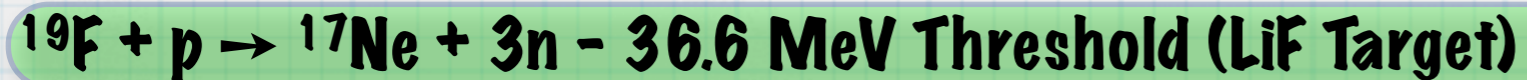
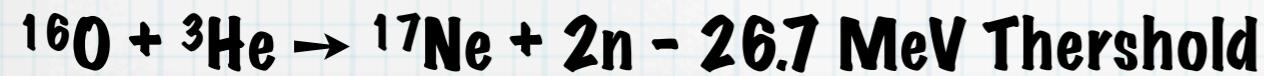


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# PROSPECTS FOR PRODUCTION IN ISRAEL

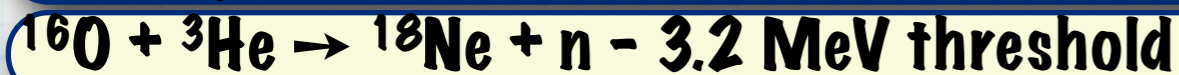
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## <sup>18</sup>Ne: Explored as a possible source for $\beta$ beams



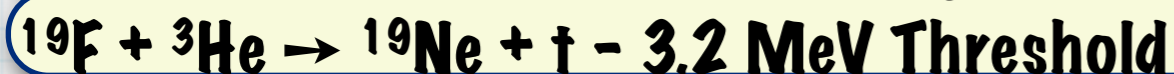
- 13 MeV Protons of LiF target,  $3 \times 10^{-5}/\text{p}$  (Takayama (2009))



## <sup>19</sup>Ne:



- 4 MeV Threshold
- ( $10^{-3}/\text{p}$  for thick LiF, Kitwanga et. al (1990)).



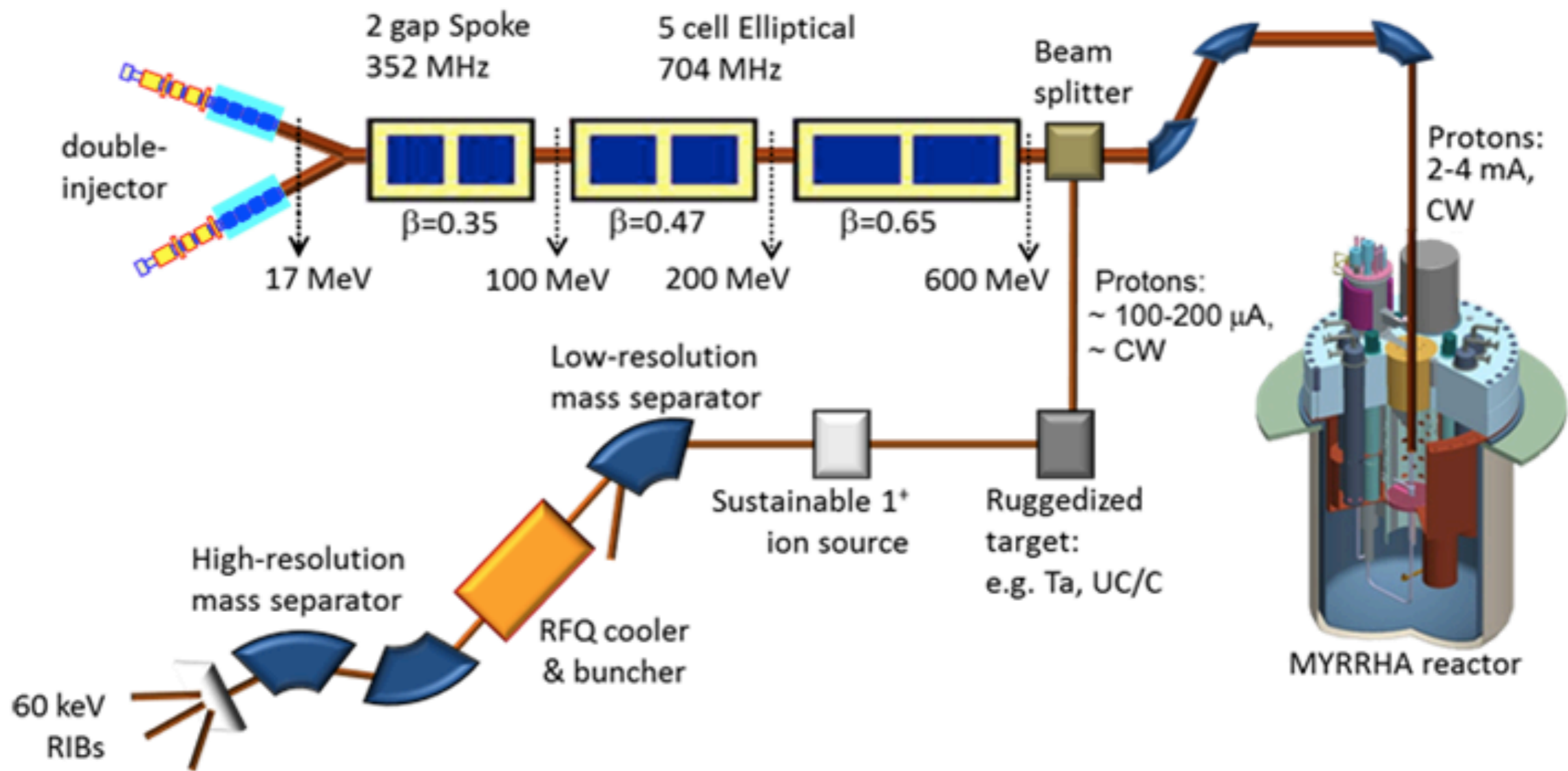
SARAF

VDC

NG

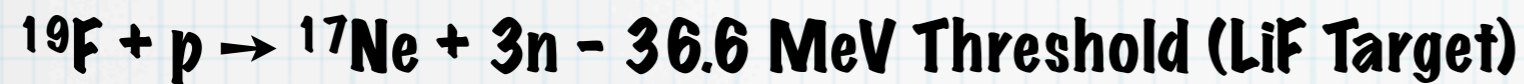
SARAF (Maybe)





# POSSIBLE PRODUCTION SCHEMES @ MYRRHA W/ PROTONS (NON-SPALLATION)

**<sup>17</sup>Ne:**



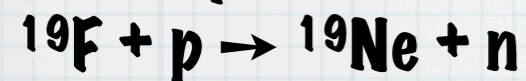
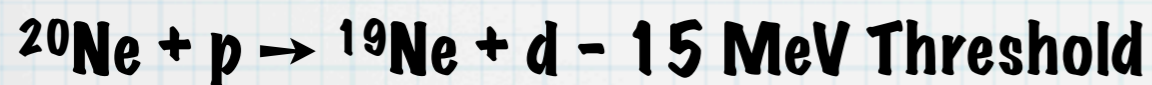
**<sup>18</sup>Ne:** Explored as a possible source for  $\beta$  beams



- 13 MeV Protons of LiF target,  $3 \times 10^{-5}/p$  (Takayama (2009))

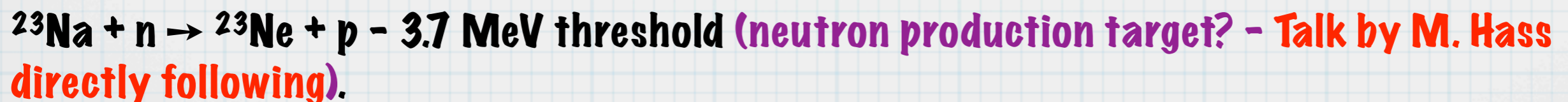


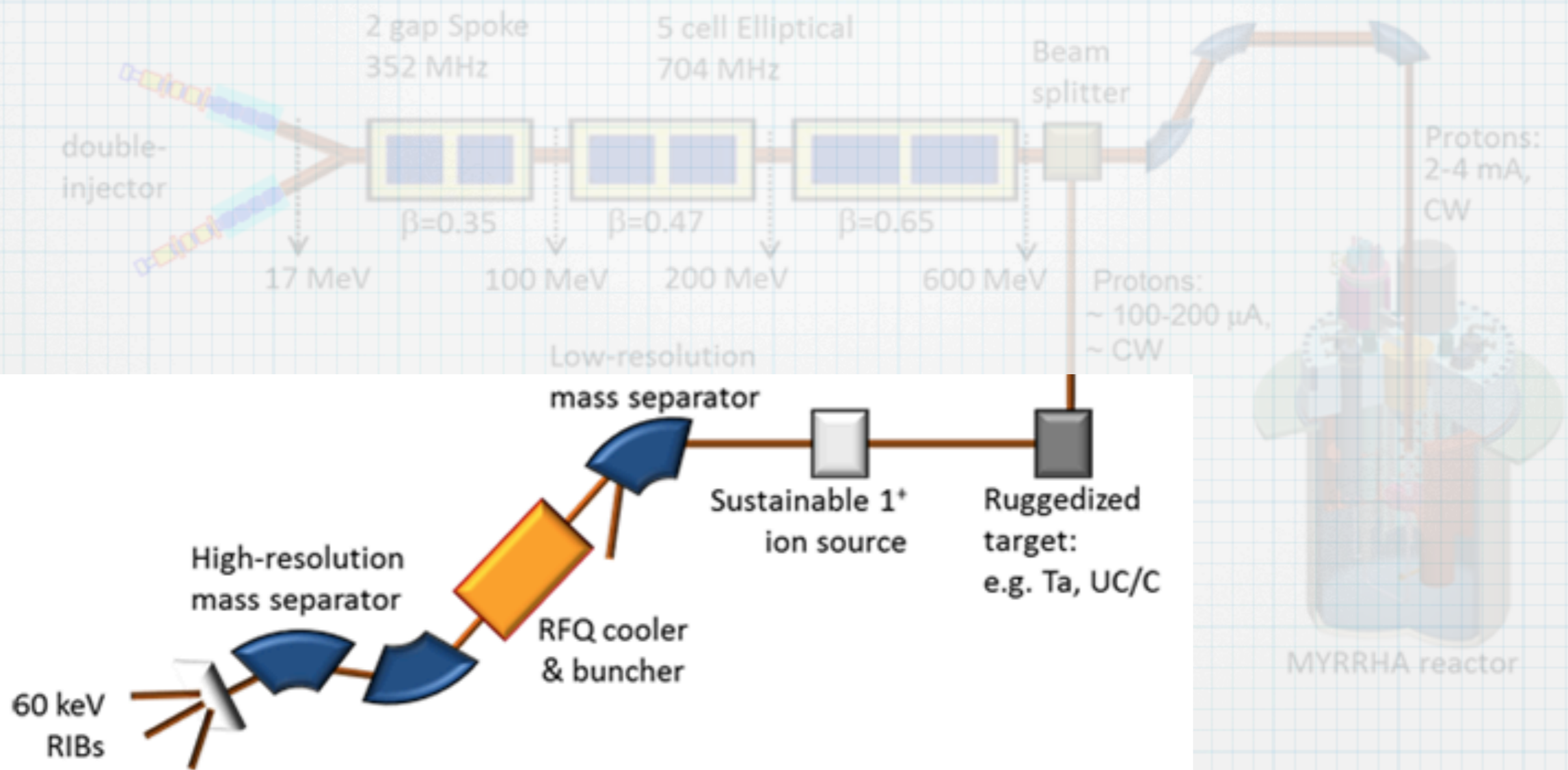
**<sup>19</sup>Ne:**



- 4 MeV Threshold
- ( $10^{-3}/p$  for thick LiF, Kitwanga et. al (1990)).

**<sup>23</sup>Ne:**







# Other things to do with neon

- Neon EIT (Already have grant for developing this).
- Atomic beam lithography.
- Autoresonance in  $\text{Ne}^*$  MOT.

# SUMMARY

- \* Neon is an interesting candidate for beyond SM search (but this of course was known long ago).
- \* HUJI has an active program which is aimed at achieving accurate measurements for Ne isotope.
- \* Production schemes are possible in Israel (at least some of them are) and are also certainly possible @ MYRRHA.
- \* We would welcome the opportunity to run @ MYRRHA in a few years.